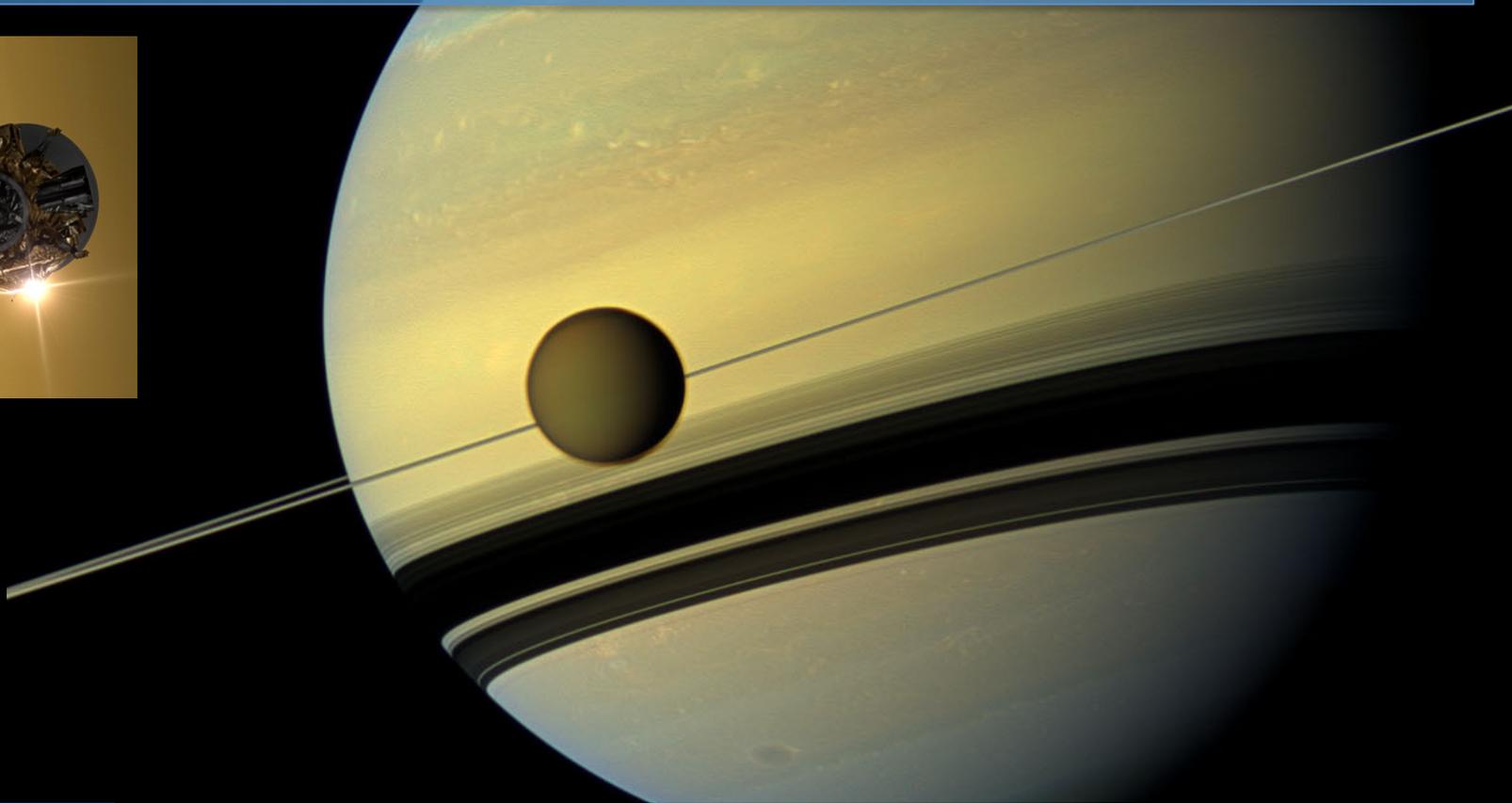
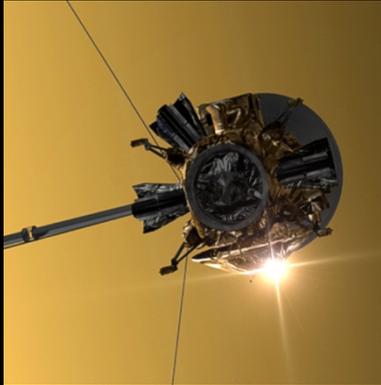


# SATURN'S SEASONALLY CHANGING ATMOSPHERE: THERMAL STRUCTURE, COMPOSITION, AND CHEMISTRY [CHAPTER TEN]



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OXFORD

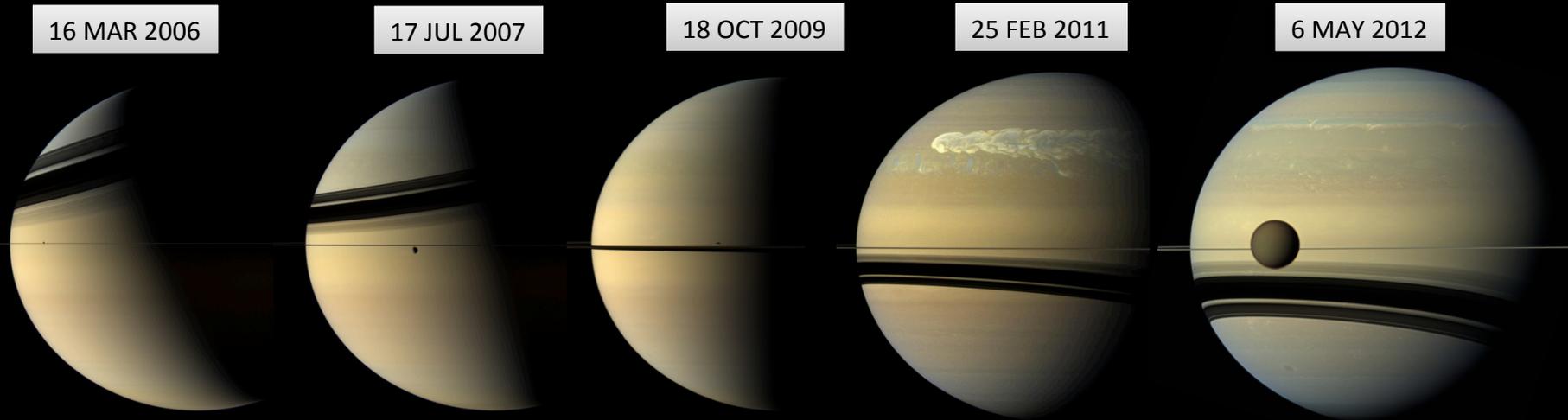


**Leigh N. Fletcher, J. Moses, T.K.  
Greathouse, S. Guerlet, R. West**



THE ROYAL  
SOCIETY

# Seasonal Saturn



- Saturn has orbited the Sun 13.5 times since Galileo's first observations of the 'strange appendages'.
- Only for the last 1.5 have we been able to study seasonal thermal changes.
- ...and only for the last 0.33 have we been studying from Saturn orbit.

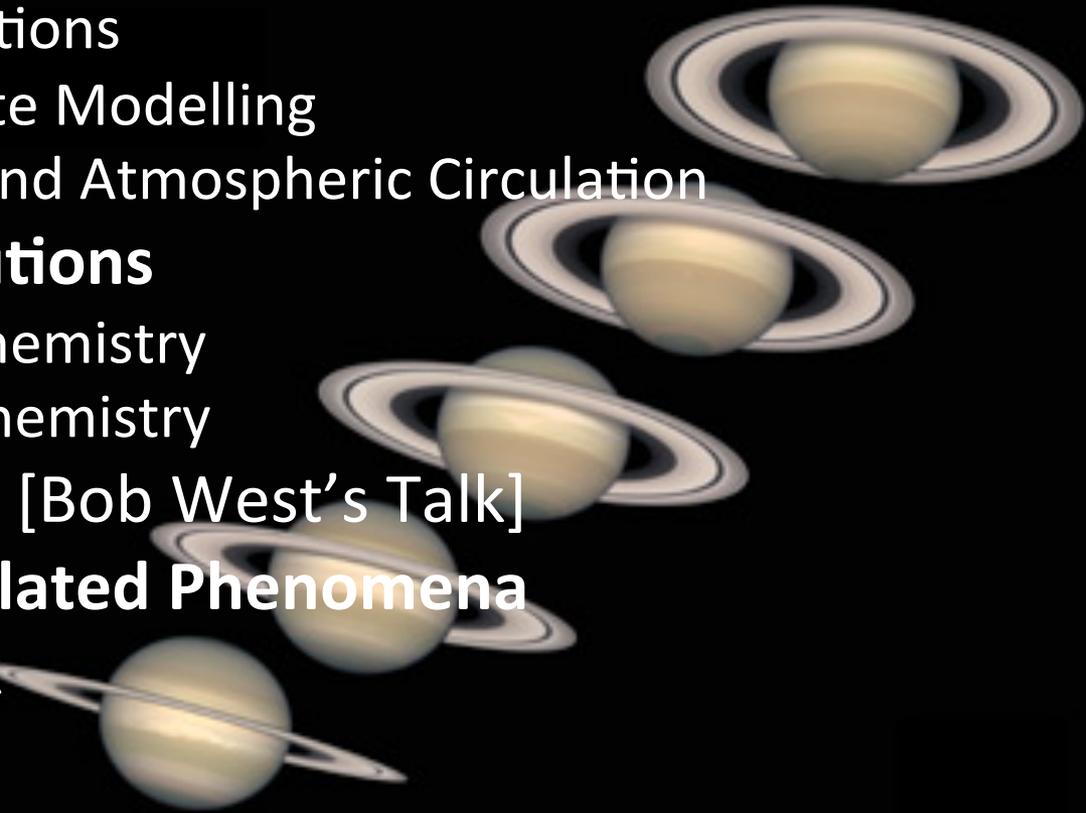
# Chapter Ten Outline

PART ONE

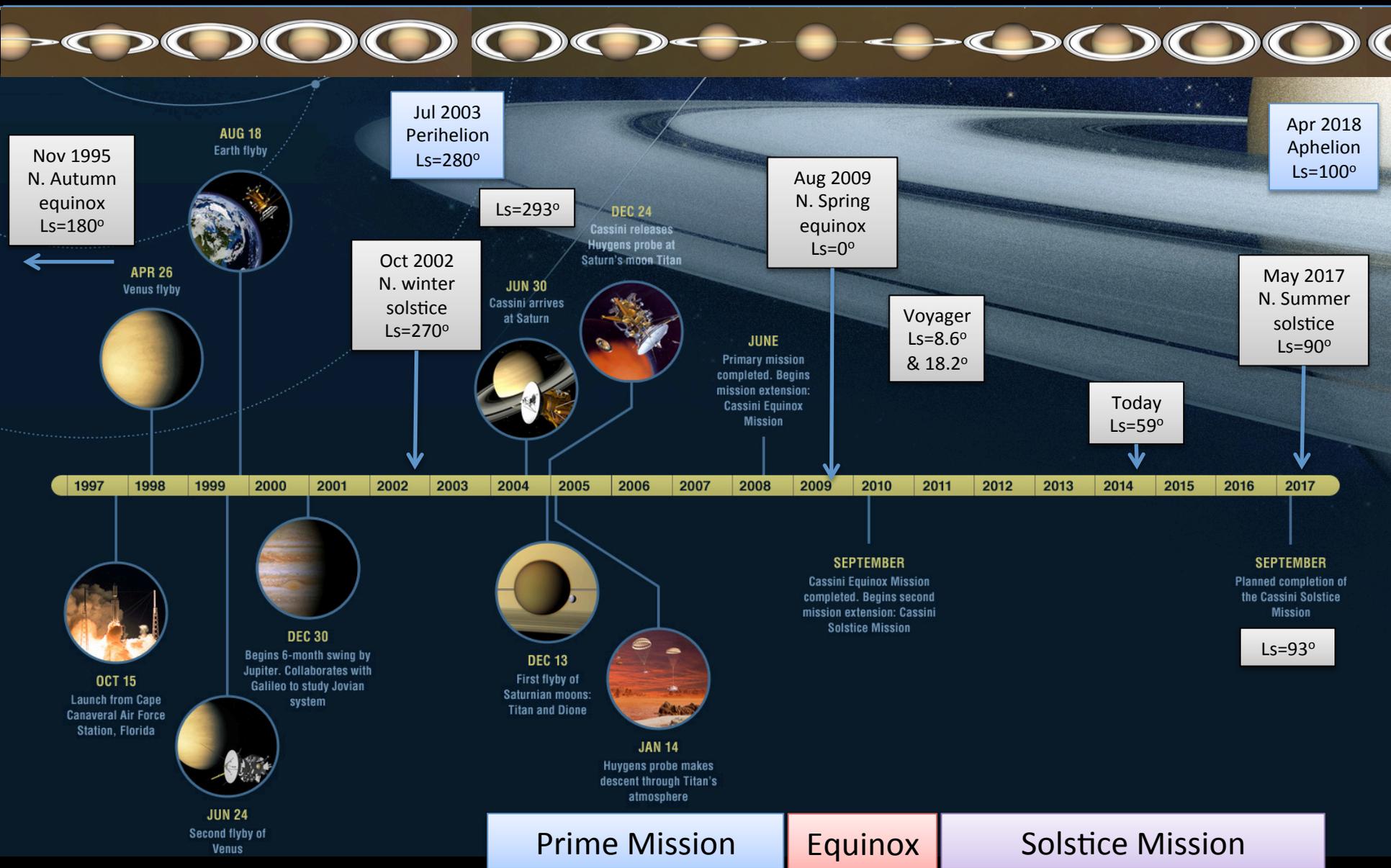
1. Introduction
2. **Seasonally-Evolving Temperature Structure**
  - a) Cassini Observations
  - b) Radiative Climate Modelling
  - c) Temperatures and Atmospheric Circulation

PART TWO

3. **Chemical Distributions**
  - a) Tropospheric Chemistry
  - b) Stratospheric Chemistry
4. **Clouds and Hazes [Bob West's Talk]**
5. **Seasonally-Modulated Phenomena**
6. **Seasonal Outlook**



# Cassini's Exploration of Seasons



Prime Mission      Equinox      Solstice Mission

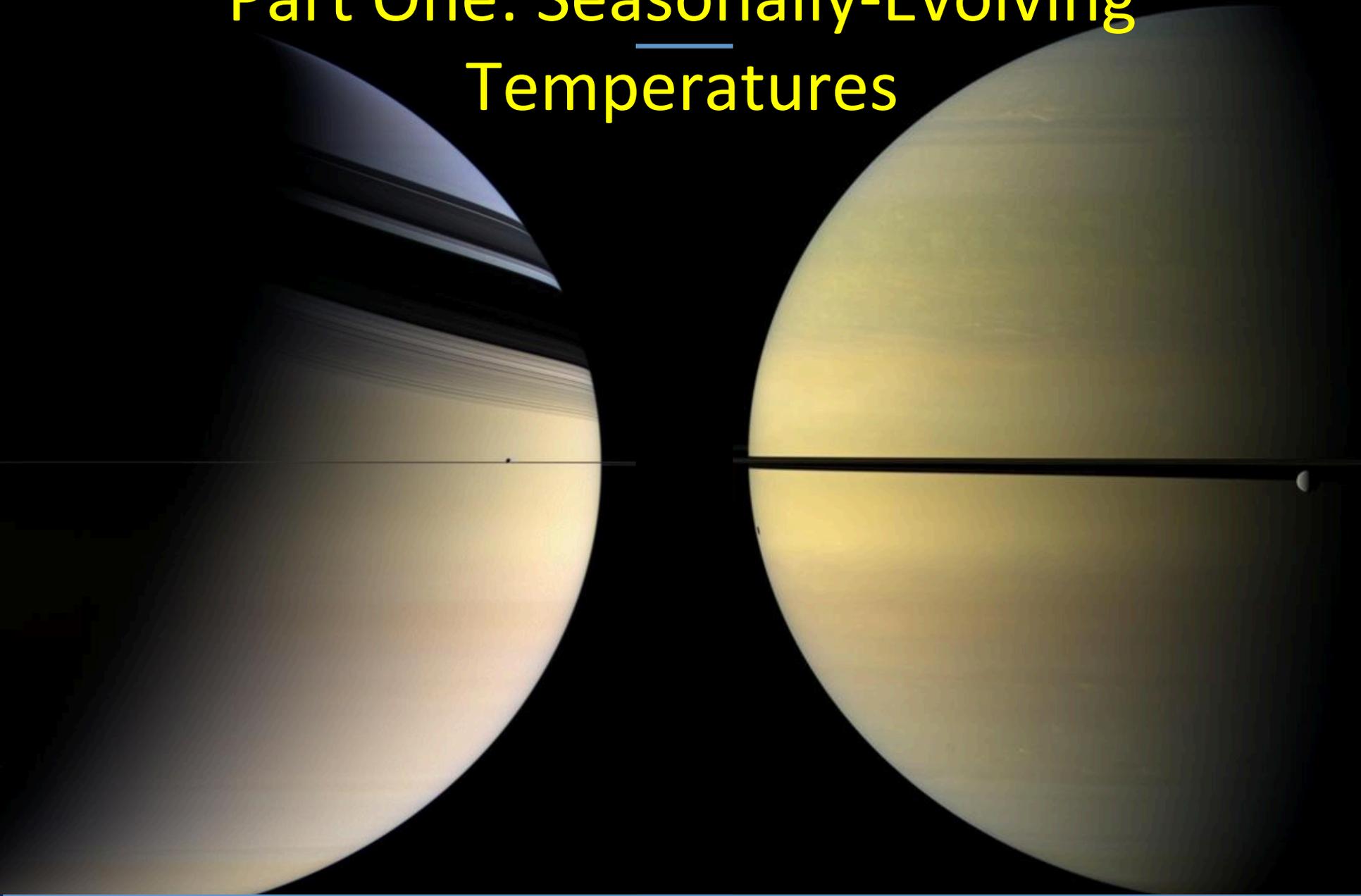
# What do we Expect?

- **Warm summers and cold winters** with phase lags that increase with depth
  - Temperature asymmetries, variable windshears.
  - Warmer summers in the south (nearer perihelion) than the north.
- **Atmospheric circulation** to redistribute energy.
  - Localized dynamics perturbing hemispheric asymmetries; large-scale overturning.
- **Variable ionization/photochemistry rates** influencing population of stratospheric & tropospheric species/hazes.
  - Chemical and aerosol hemispheric asymmetries.
- **Varying stratification & wave transmissivity**
  - Altering propagation of dynamic activity (e.g., storm plumes, waves).



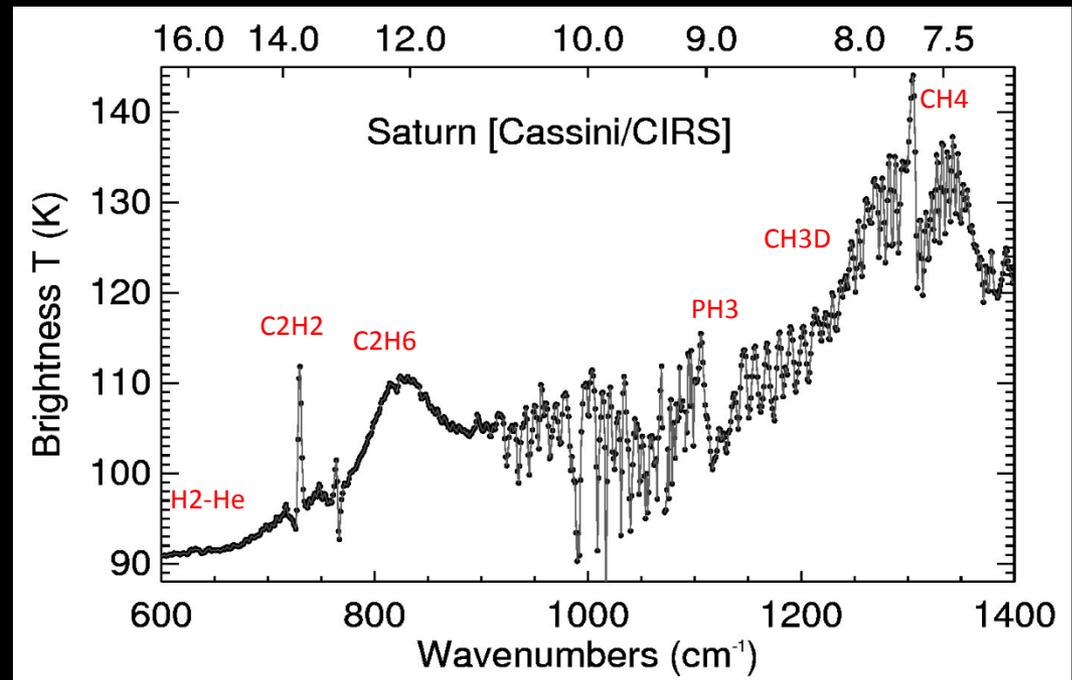
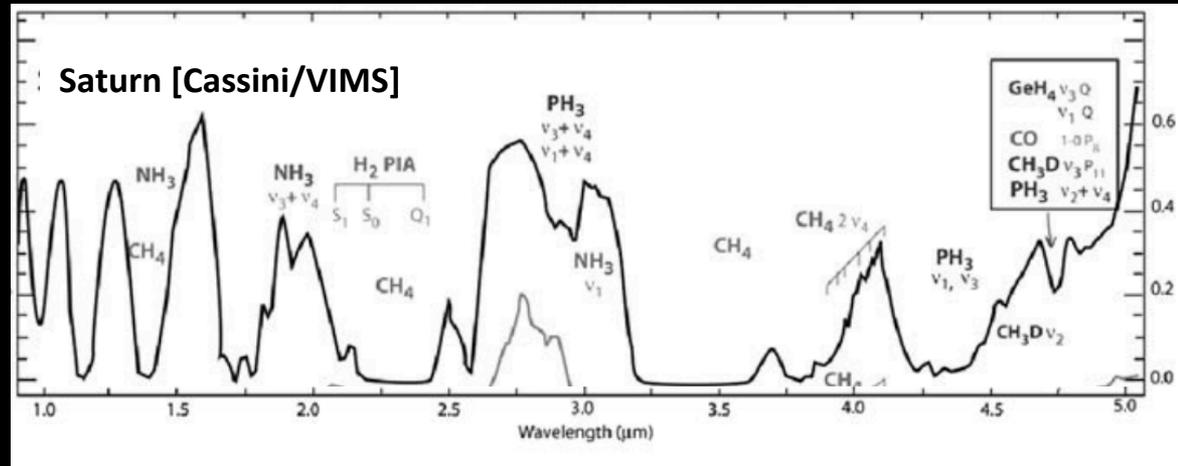
Long term remote sensing record allows us to compare hypotheses to observations

# Part One: Seasonally-Evolving Temperatures



# Saturn's Infrared Spectrum

- **Remote sensing toolkit for seasonal studies:**
  - Reflected sunlight for clouds/hazes/composition.
  - Thermal emission for  $T(p)$  and composition.
- Reflectivity changes (UV-Vis) to be discussed in Bob West's talk.
- Temperature/composition variations covered here.



# Detection of Asymmetries: 1970s & Voyager

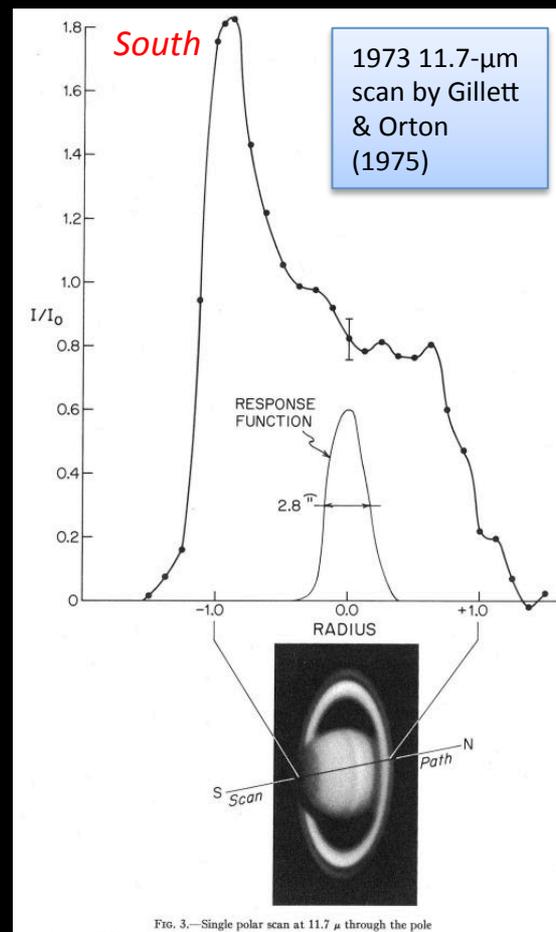
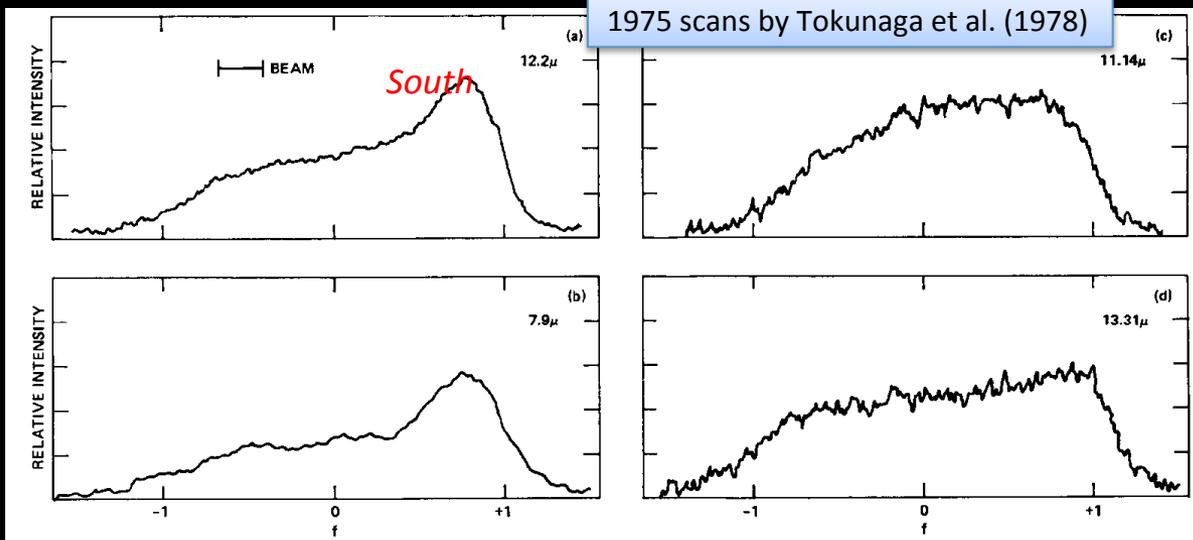
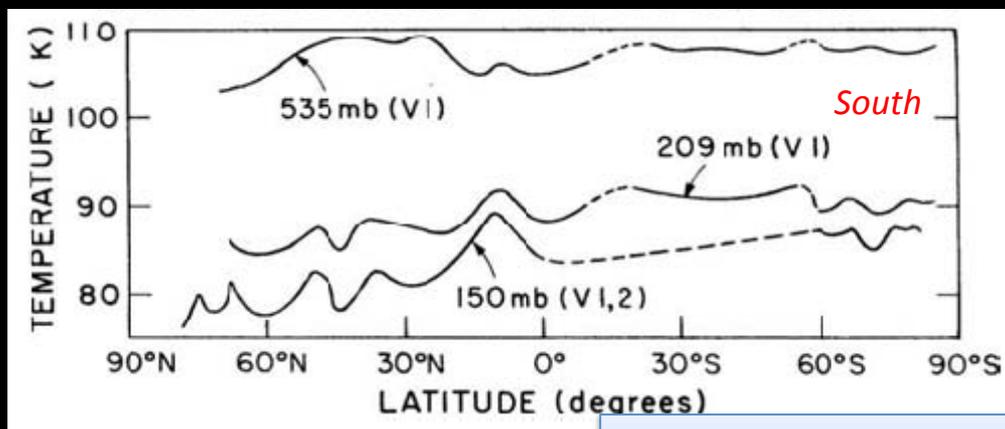


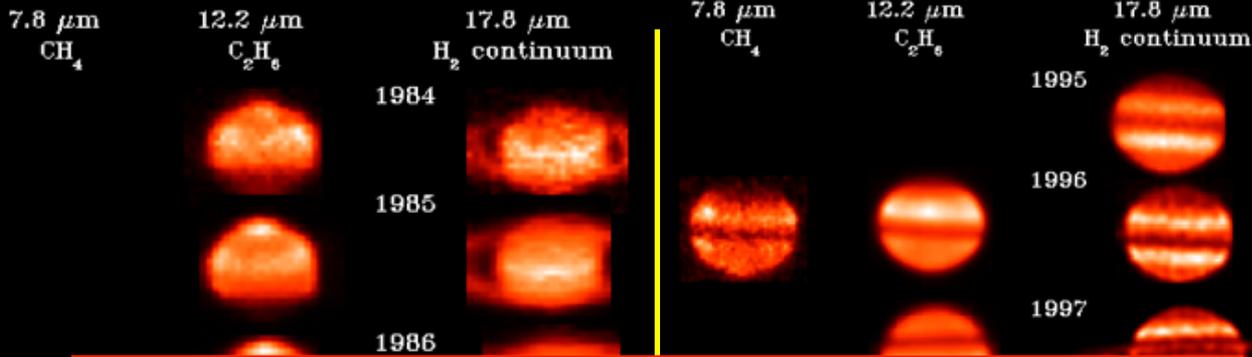
FIG. 3.—Single polar scan at 11.7  $\mu$  through the pole



1980/81 Voyager/IRIS observations (Hanel et al., 1981; Conrath & Pirraglia (1983))

- Detection of asymmetries in troposphere and stratosphere in southern summer/autumn.

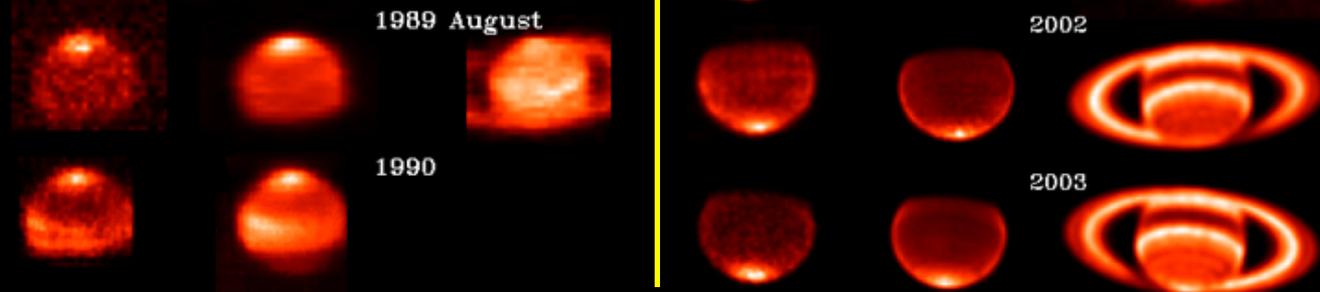
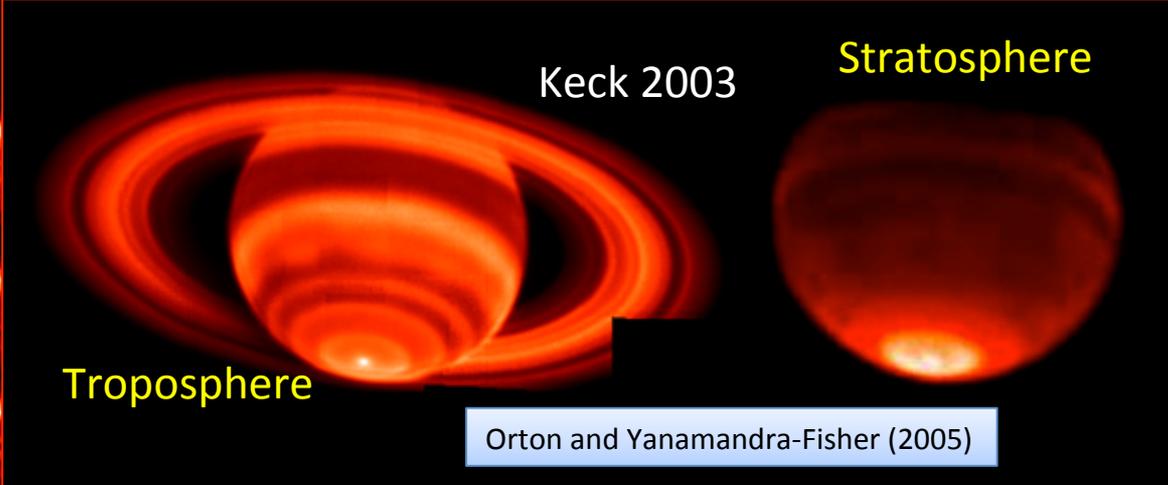
# 2D Thermal Imaging: From Voyager to Cassini



Northern spring equinox, March 1980  
Ls=0

Northern autumnal equinox, Nov 1995  
Ls=180

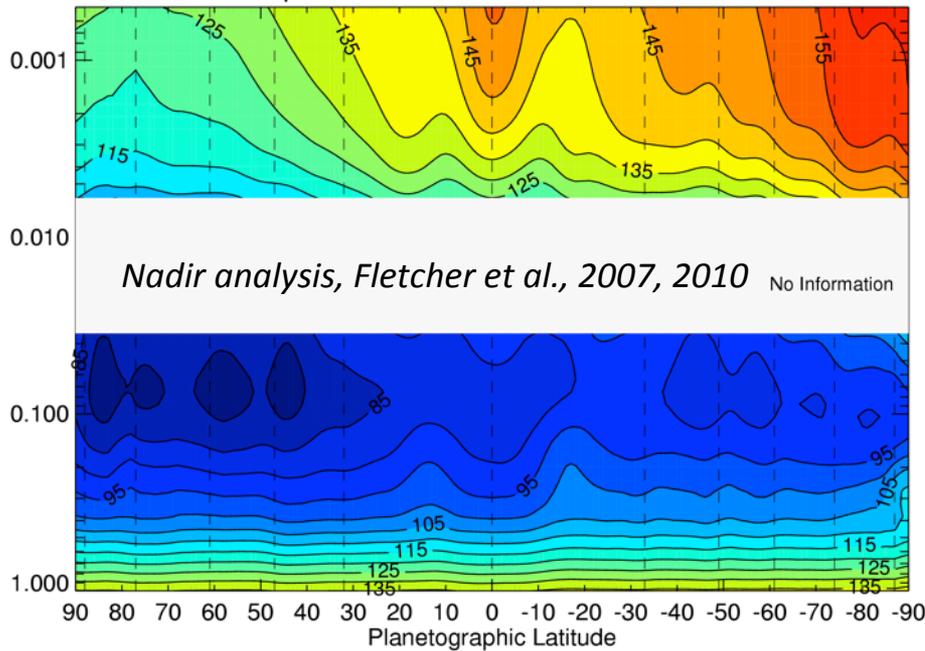
Northern summer solstice, Dec 1987  
Ls=90



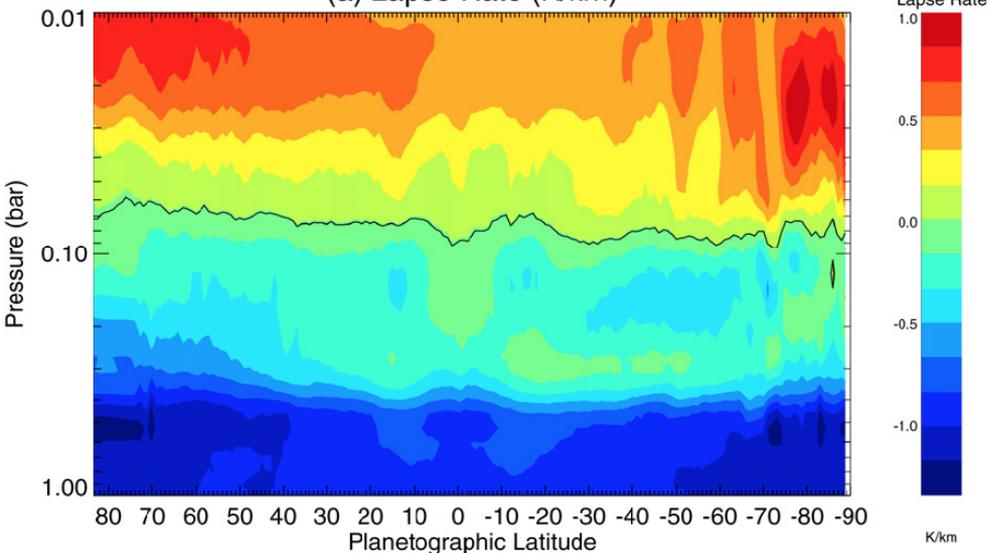
Northern winter solstice, Oct 2002  
Ls=270

# Summertime Asymmetries

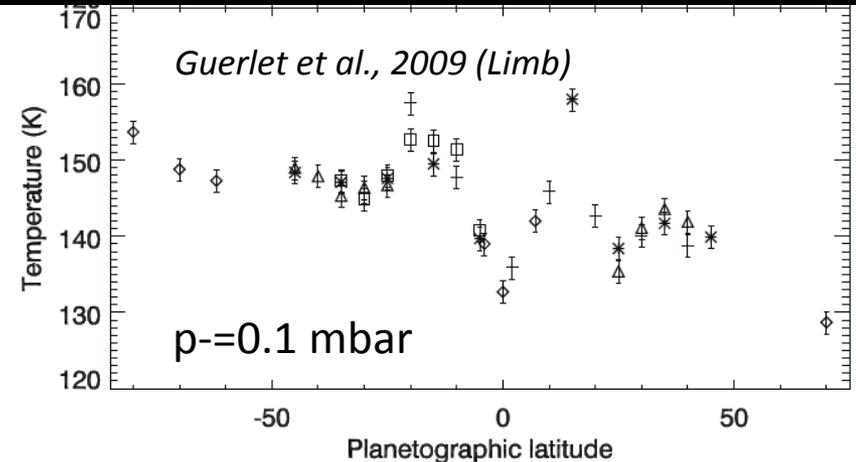
Temperature from Cassini Prime Mission



(a) Lapse Rate (K/km)

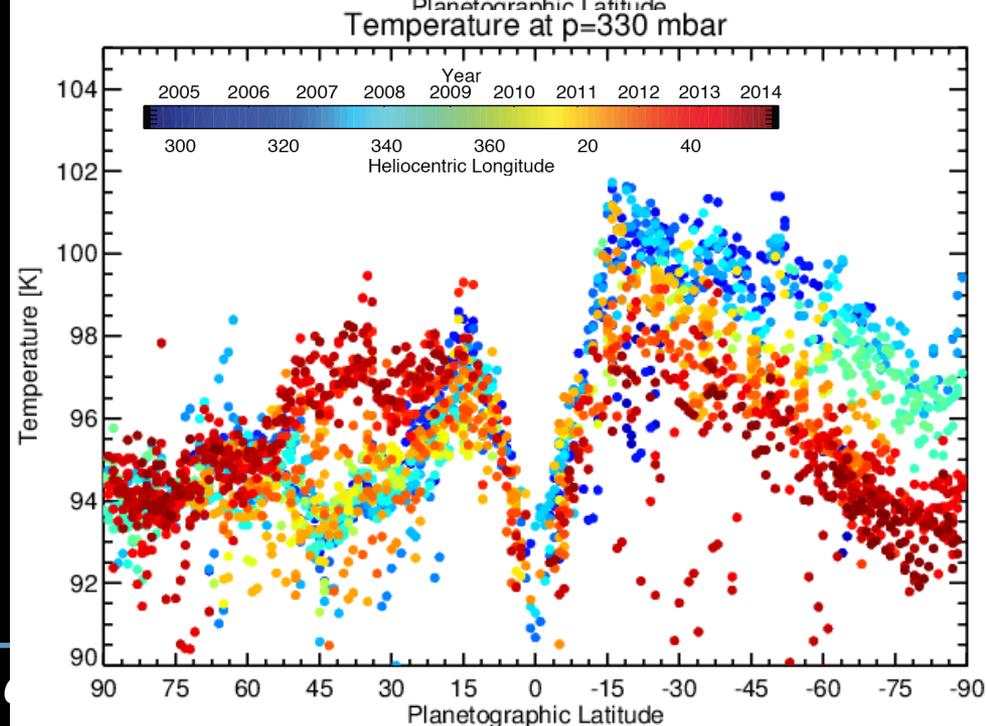
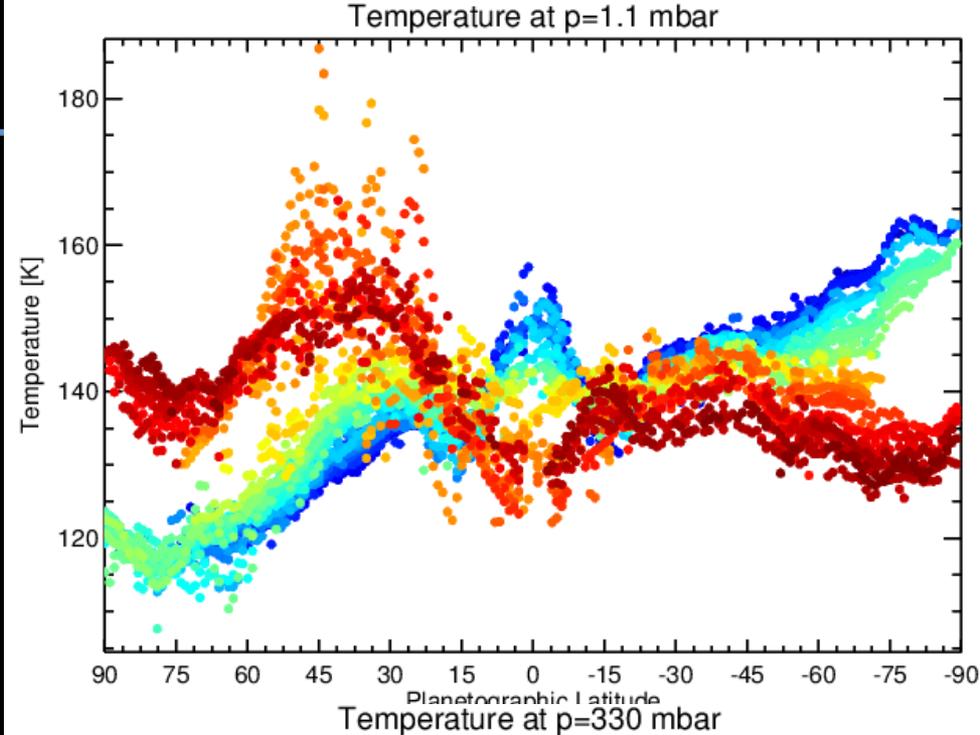


- Stratospheric 40 K asymmetry; 10 K tropopause asymmetry; no asymmetry  $p > 500$  mbar.
- Belt/zone contrasts superimposed onto hemispheric asymmetries.
- 80-mbar tropopause and 350-500 mbar RC boundary both slightly deeper in summer hemisphere.



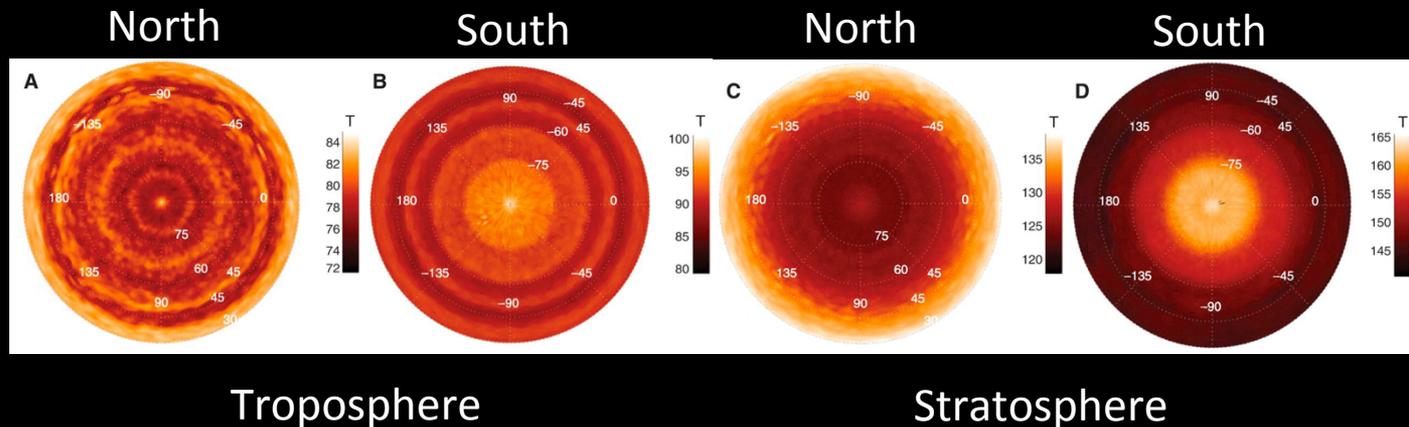
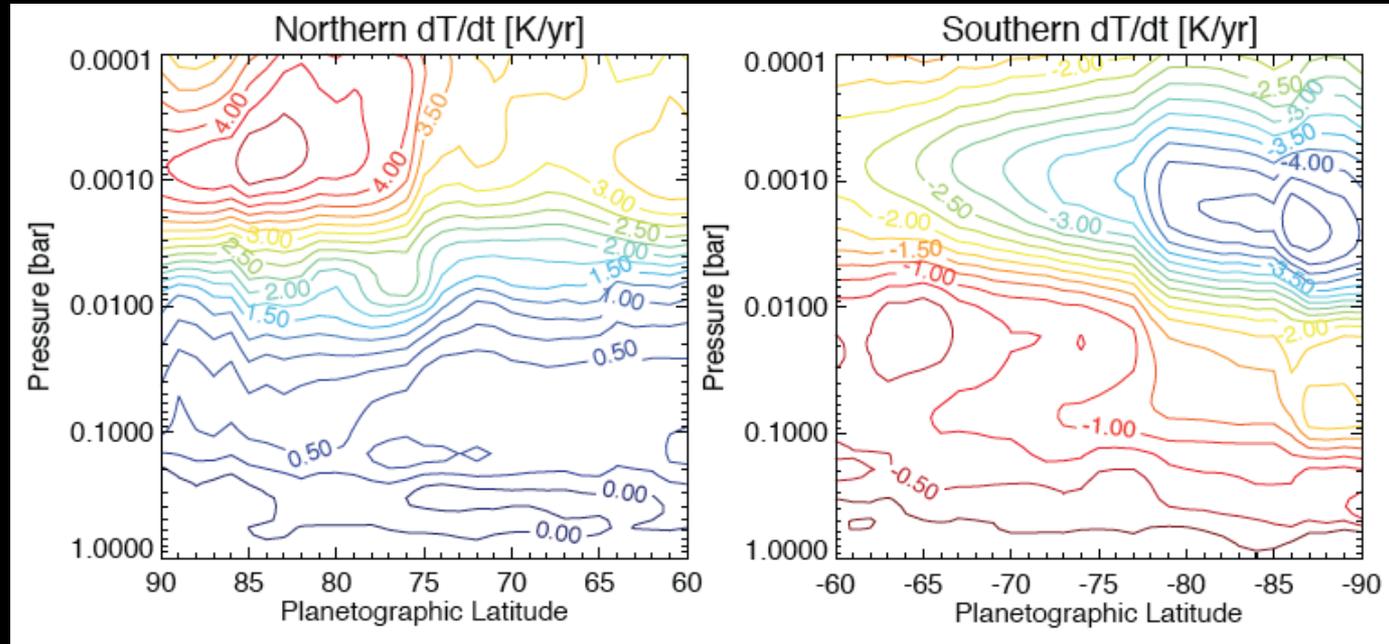
# Temperature Evolution

- Nadir temperature sounding for a 1/3 Saturn year
  - Updated from Fletcher et al. (2010) pre-equinox observations.
- **Thermal asymmetries reversing.**
  - Northern mid-latitudes and poles are changing the fastest.
  - Equatorial oscillation associated with SSAO.
- Trends governed by radiative heating/cooling & adiabatic compression/expansion.
- Implications for:
  - Thermal wind balance.
  - Saturn emitted power (Li et al., 2010).



# Polar Seasonal Change

- Polar vortices – most dramatic manifestation of seasonal effects.
- Closer inspection reveals most rapid  $dT/dt$  (and  $dq/dt$ ) within the polar vortices.
- Either:
  - vigorous vertical motion and adiabatic heating/cooling, or
  - aerosol contribution to radiative budget.



Troposphere

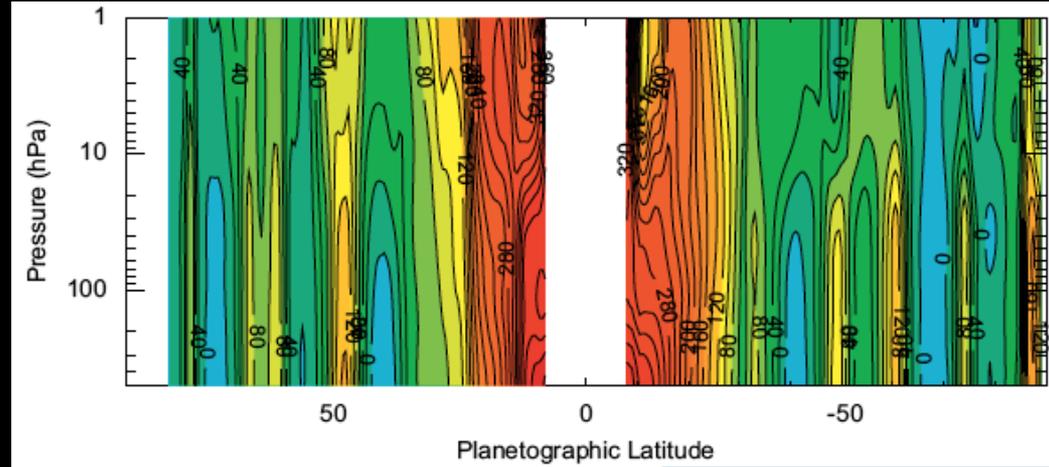
Stratosphere

# Implications for Wind Shear

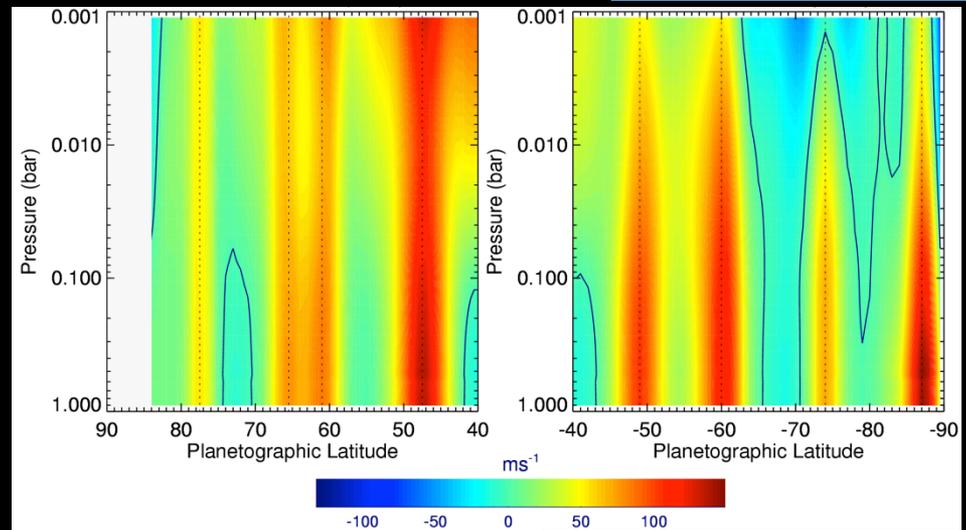
$$f \frac{\partial u}{\partial \ln(p)} = \frac{R}{a} \frac{\partial T}{\partial \psi} = R \frac{\partial T}{\partial y}$$

$$f \frac{\partial v}{\partial \ln(p)} = -\frac{R}{a \cos \psi} \frac{\partial T}{\partial \phi} = -R \frac{\partial T}{\partial x}$$

- Seasonal  $dT/dy$  strongly affects vertical shear.
- Decay with altitude stronger in the summer hemisphere; increasingly barotropic (height-independent) zonal jet structure in the winter hemisphere.
- Influences stratification/trapping vertically-propagating waves.

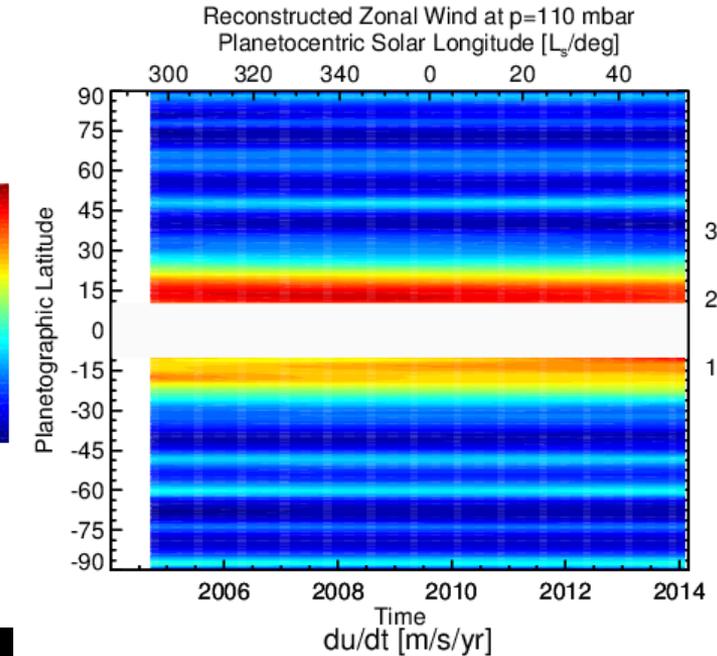
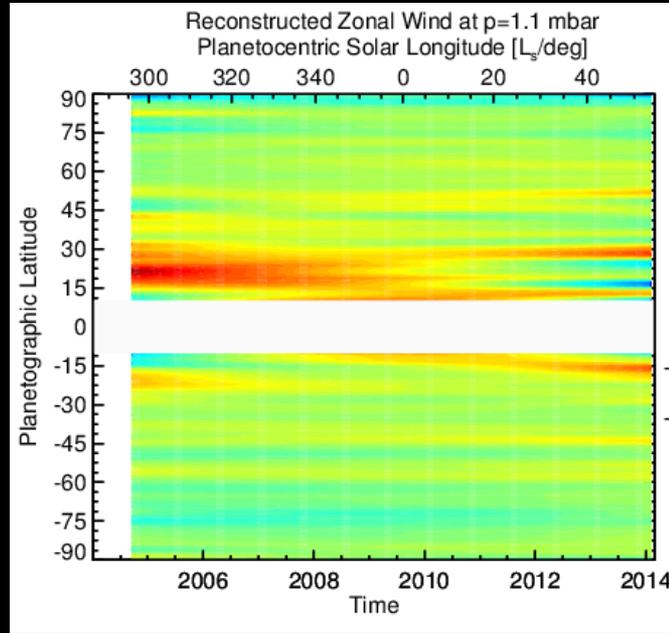


*Read et al. (2009)*

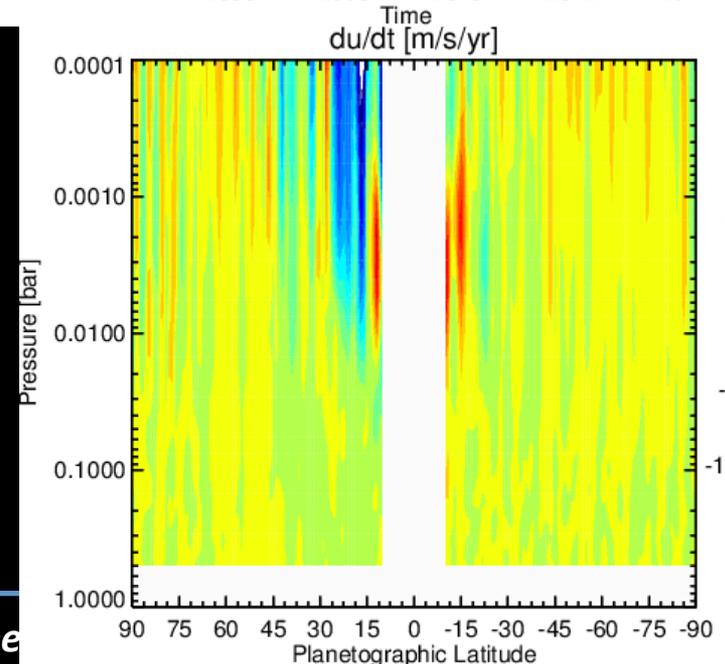


*Fletcher et al., (2008)*

# Evolving Stratospheric Winds

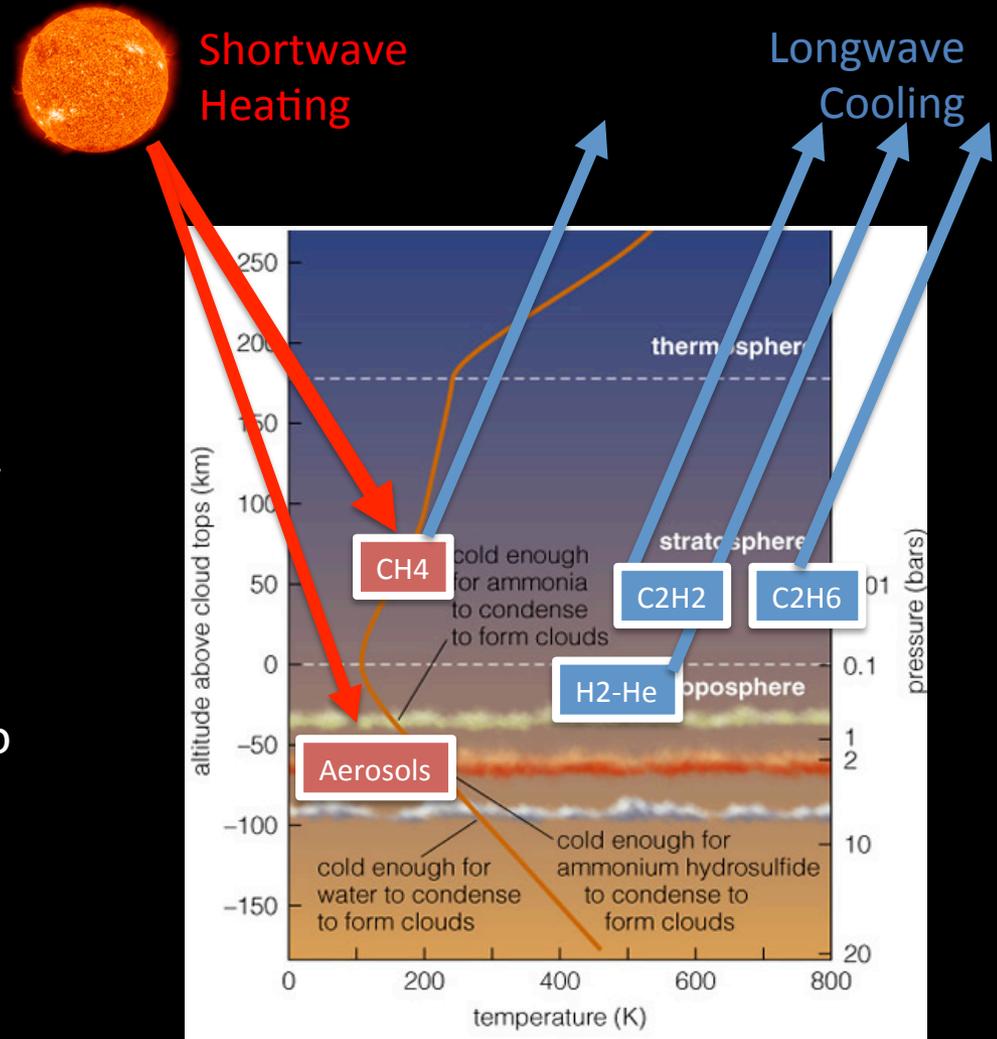


- Zonal winds reconstructed from 10-years of  $T(p)$ .
- Seasonal  $dT/dy$  promotes positive vertical shear in the winter hemisphere and negative vertical shear in the summer hemisphere (e.g., Friedson and Moses, 2012).
- Stratosphere: Should see **northern prograde winds weakening**, **southern prograde winds strengthening** – partial confirmation by CIRS (unpublished).



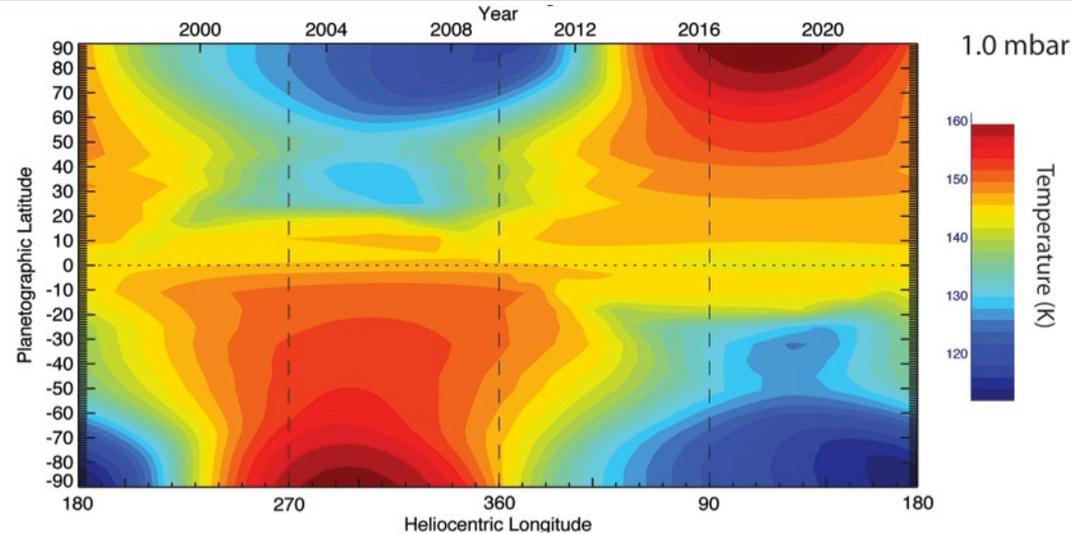
# Radiative Climate Models

- Balance atmospheric heating (**methane and aerosols**) with key coolants (**strat. ethane/acetylene; trop H<sub>2</sub>-He**).
- Earliest 1D radiative convective models (1970s) explained temperature inversion
  - Methane VNIR absorption and UV aerosol absorption (Caldwell, 1977; Tokunaga & Cess, 1977).
- 2D (zonal and seasonal) models started in the stratosphere (Cess and Caldwell, 1979), extended into troposphere (Bezard 1984; Bezard & Gautier, 1985; Conrath et al., 1990, Barnet et al., 1992).
- Next generation of models developed for Cassini analysis.



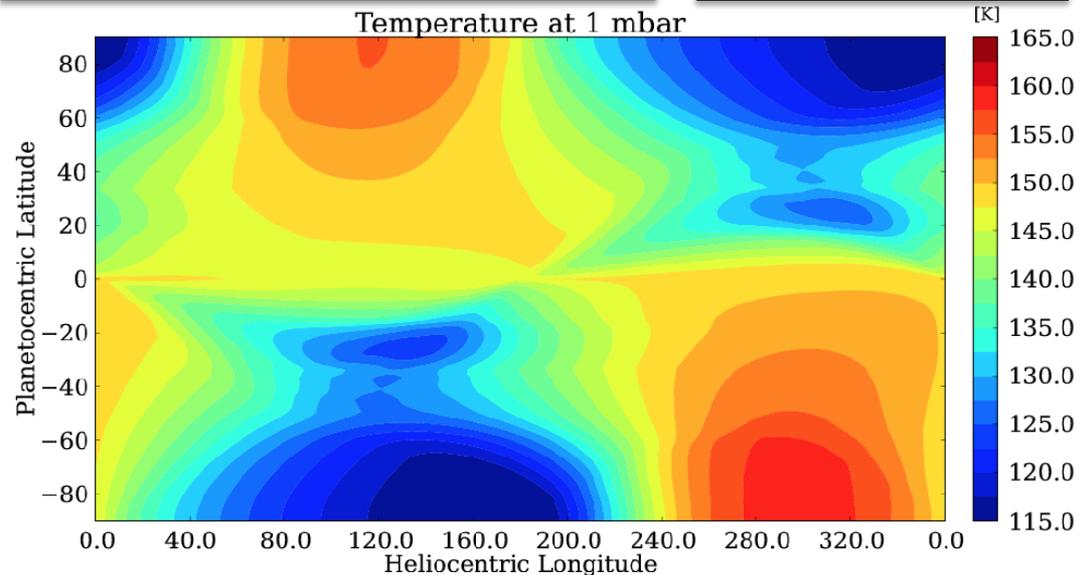
# Radiative Climate Models II

- **Greathouse et al. (2008); Friedson and Moses (2012); Guerlet et al. (2014)** most current:
  - Oblateness, ring shadowing, eccentricity all accounted for.
  - Guerlet & Friedson extend into troposphere with aerosols.
  - Friedson includes eddy transport/advective circulation.
- Reproduce large-scale stratospheric trends.
  - reproduce lag timescales and equator-to-pole contrast amplitudes.
- Greathouse uses spatially-variable  $C_xH_y$ ; Guerlet/Friedson fix at average values.

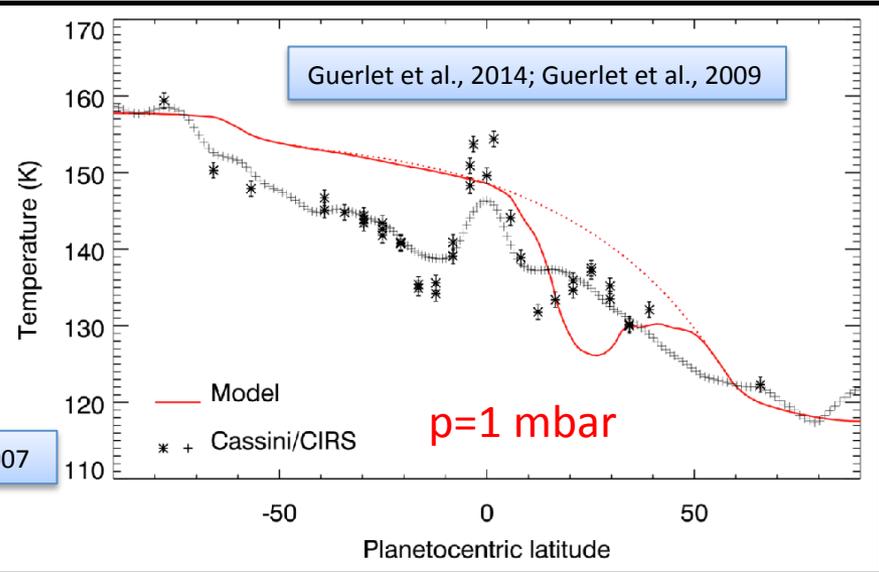
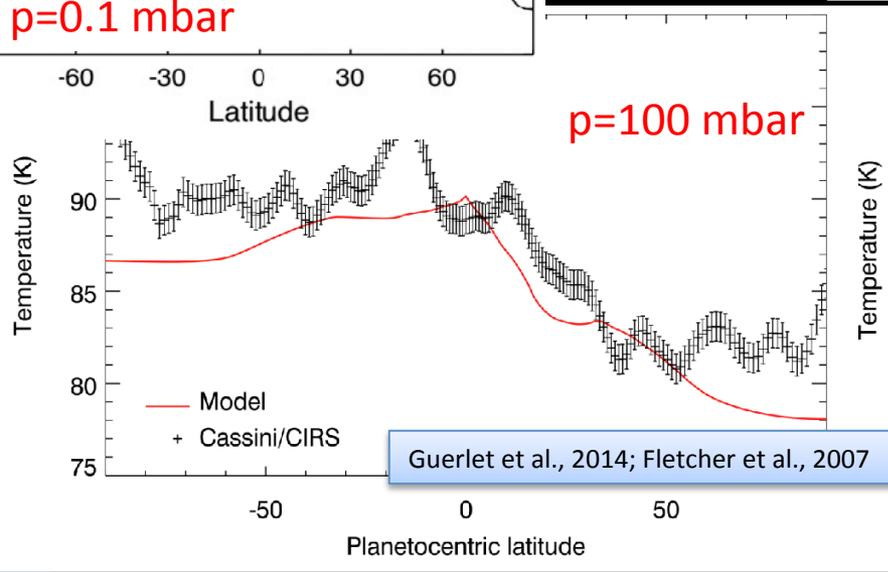
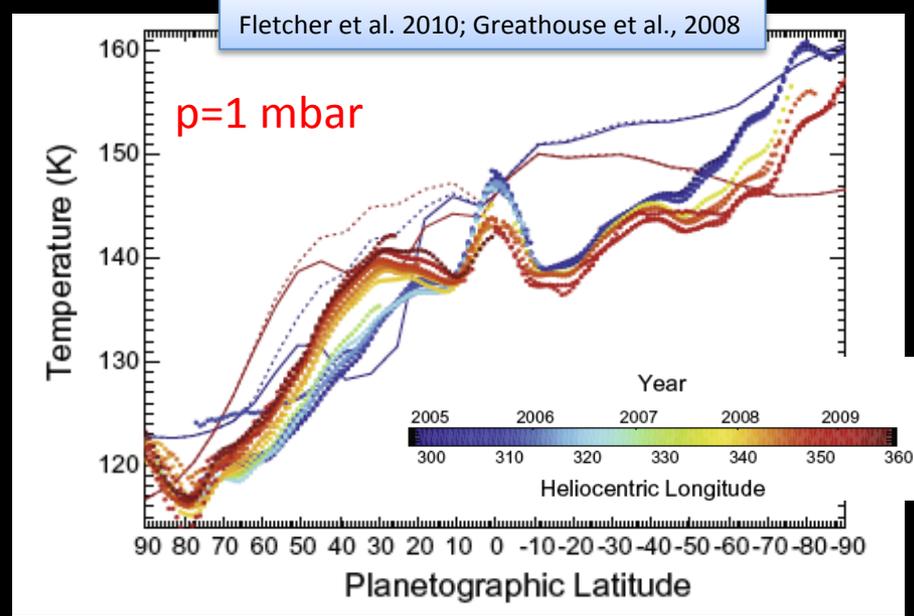
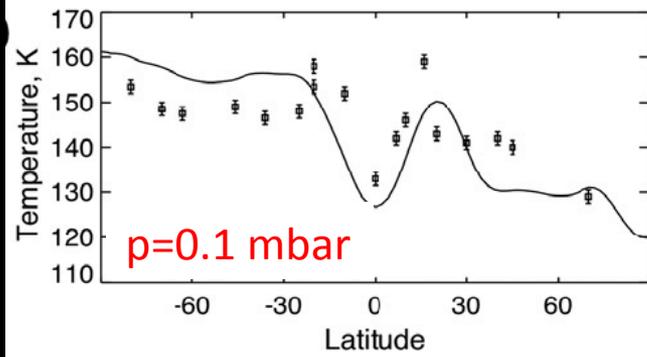
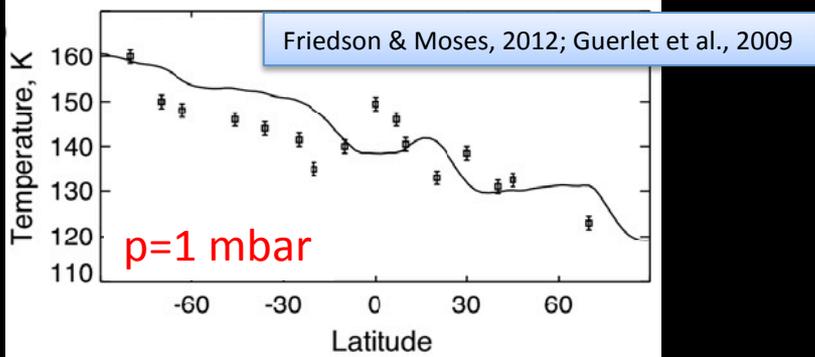


*Fletcher et al., 2010; Greathouse et al., 2008*

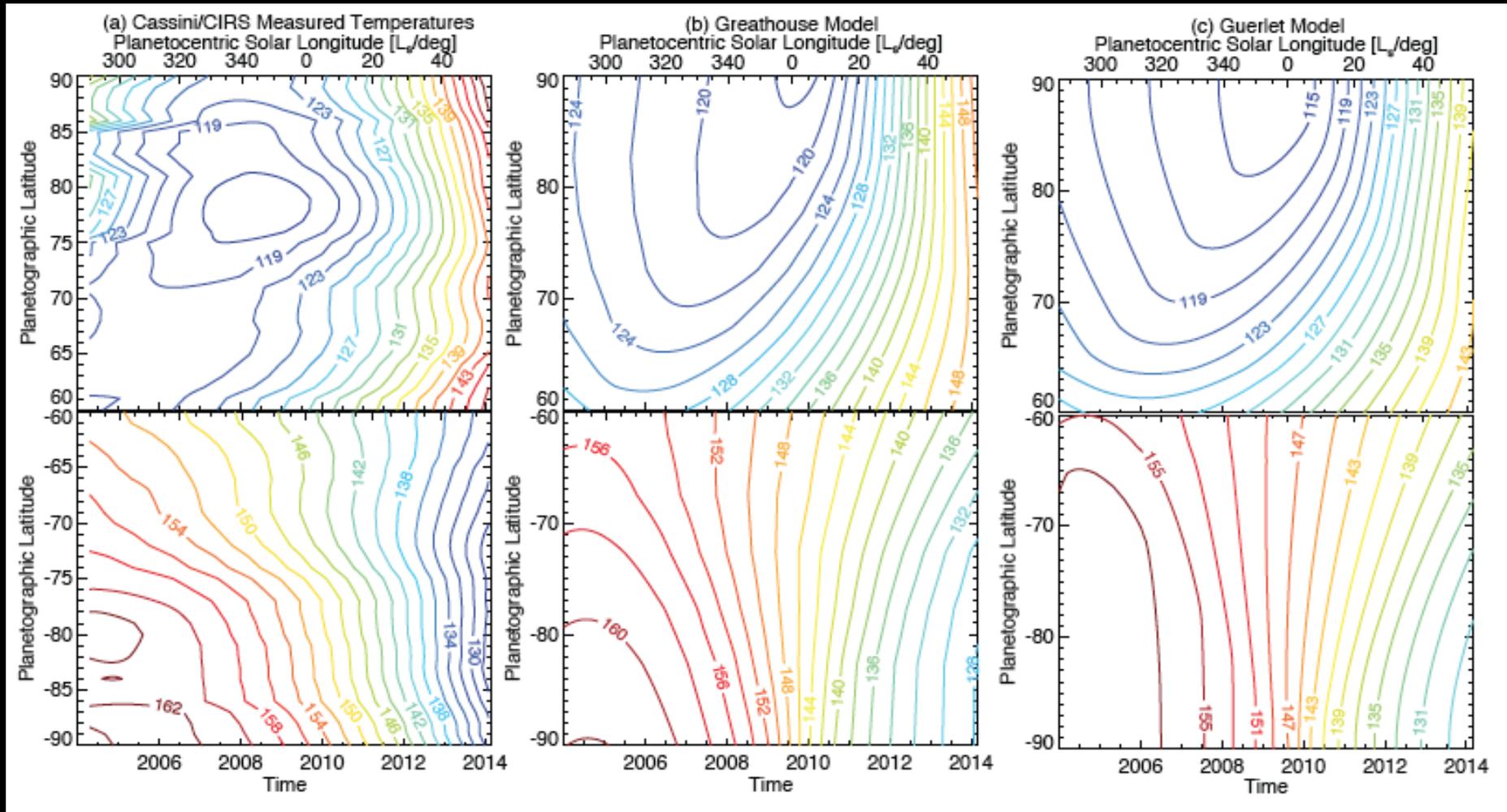
*Guerlet et al. (2014)*



# Model – Data Comparisons



# Comparison in the Polar Stratosphere

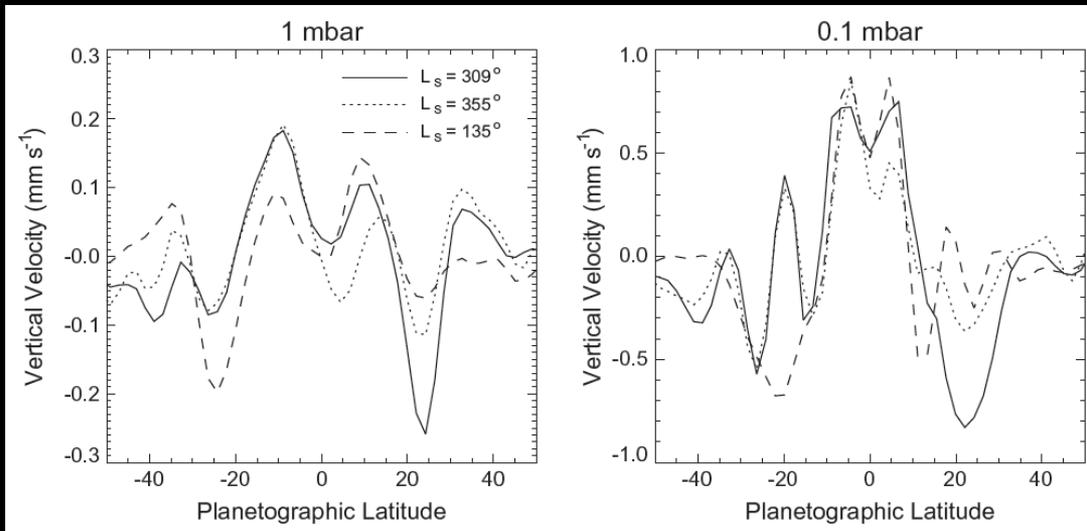


Fletcher et al., in prep.

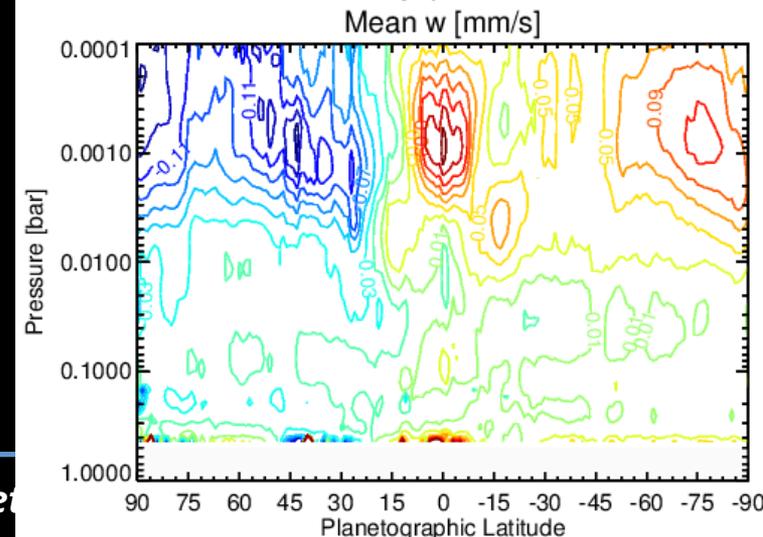
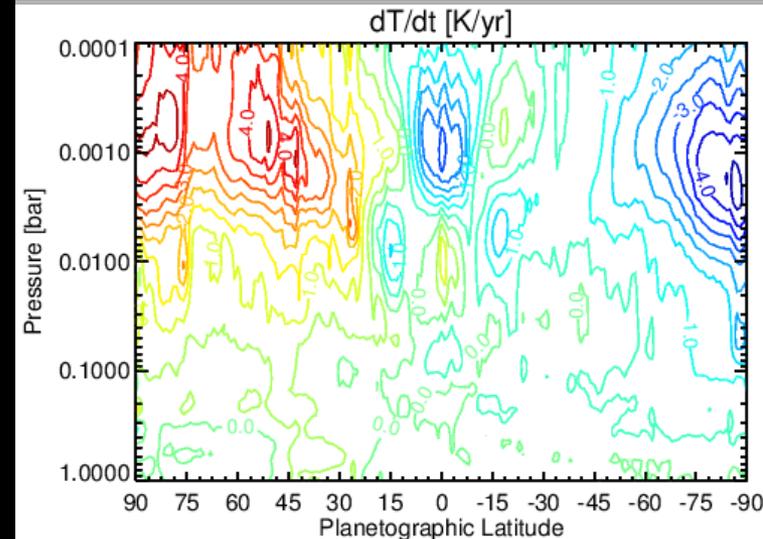
# Need for Atmospheric Circulation

- **Thermodynamic relation** of the primitive equations (e.g., Andrews, 1987).
  - Balances temporal evolution of temperature, advective heat transport and radiative relaxation to equilibrium.
- **Allows crude estimation of vertical velocities**, dominated by  $dT/dt$  term,
  - $|w| \sim 0.1$  mm/s.
- Velocities will vary seasonally –
  - Reversing Hadley circulation?
  - Doesn't capture SSAO?

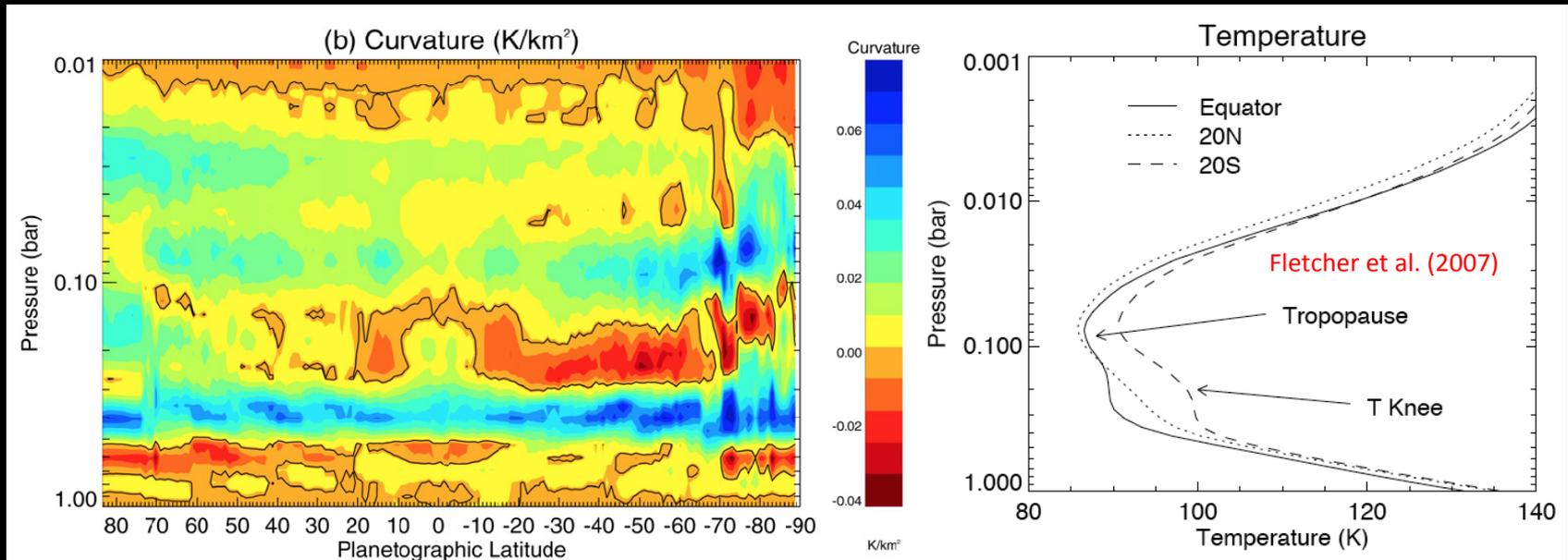
$$\left( \frac{\partial T}{\partial t} + w \left[ \frac{\partial T}{\partial z} + \frac{RT}{Hc_p} \right] \right) = \frac{Q}{\rho c_p} \approx \frac{T_E - T}{\tau}$$



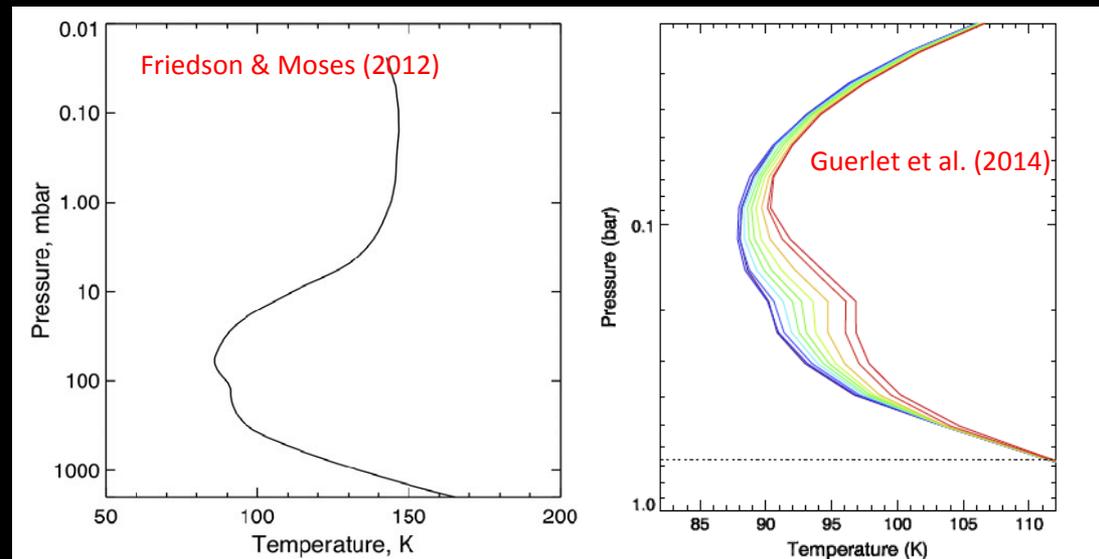
*Friedson and Moses (2012)*



# Tropospheric Temperatures

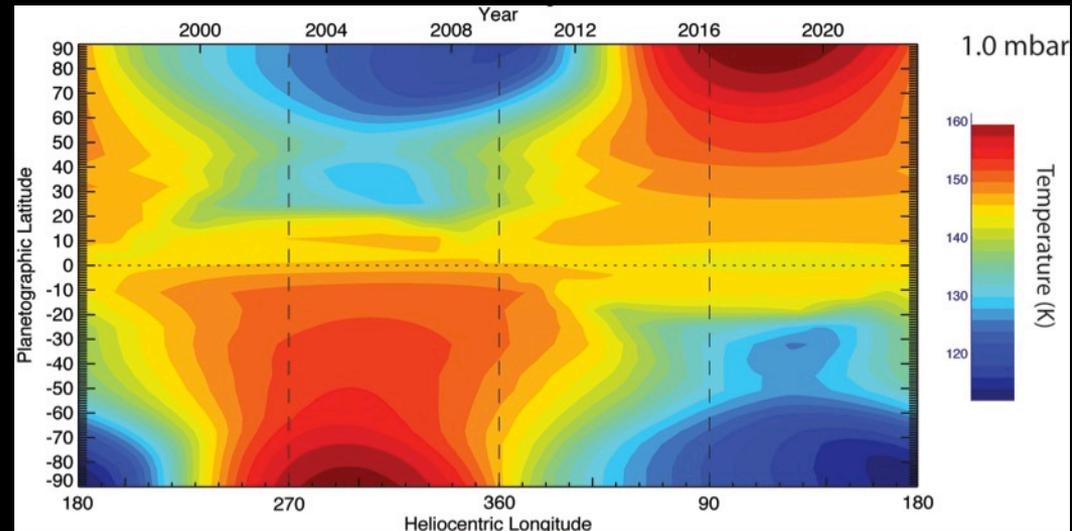


- Temperature 'knee' identified as radiative effect associated with haze (Fletcher et al., 2007).
- Hemispheric asymmetry matched VIMS 5- $\mu m$  asymmetry.
- Friedson and Moses (2012) and Guerlet et al. (2014) reproduce the temperature knee
  - Don't require aerosol asymmetry to explain effect.

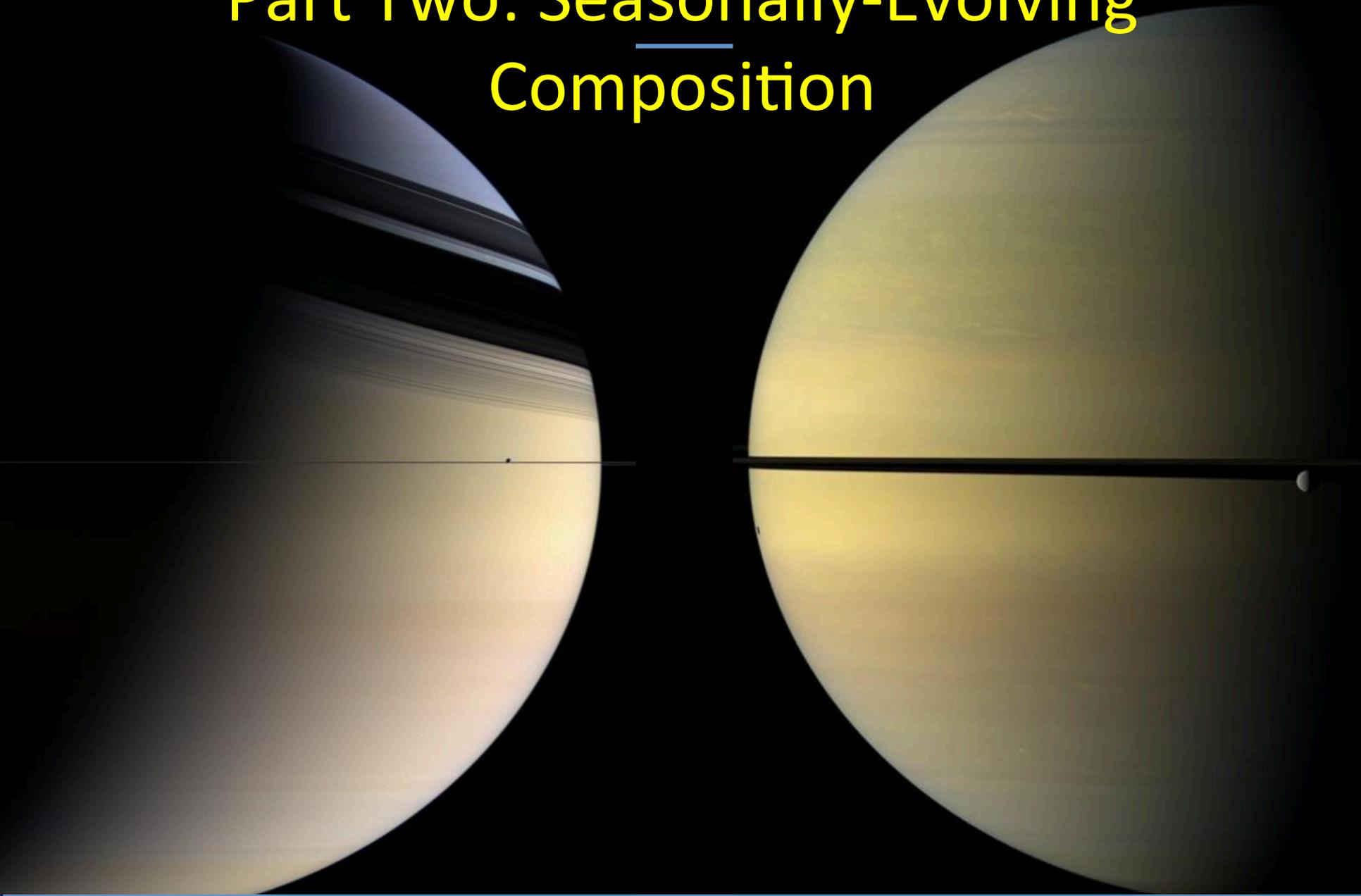


# Summary of Part One: Temperatures

1. Ground-based and Cassini observations **reveal hemispheric thermal contrasts superimposed onto belt/zone contrasts.**
2. Seasonal amplitudes and increasing phase lag with depth **reproduced by radiative climate models.**
3. **Contributions from aerosols** suggested in tropospheric 'knee' and polar vortices.
4. Implications for Saturn's **energy balance**, stratospheric **circulation**, middle-atmospheric **zonal winds.**

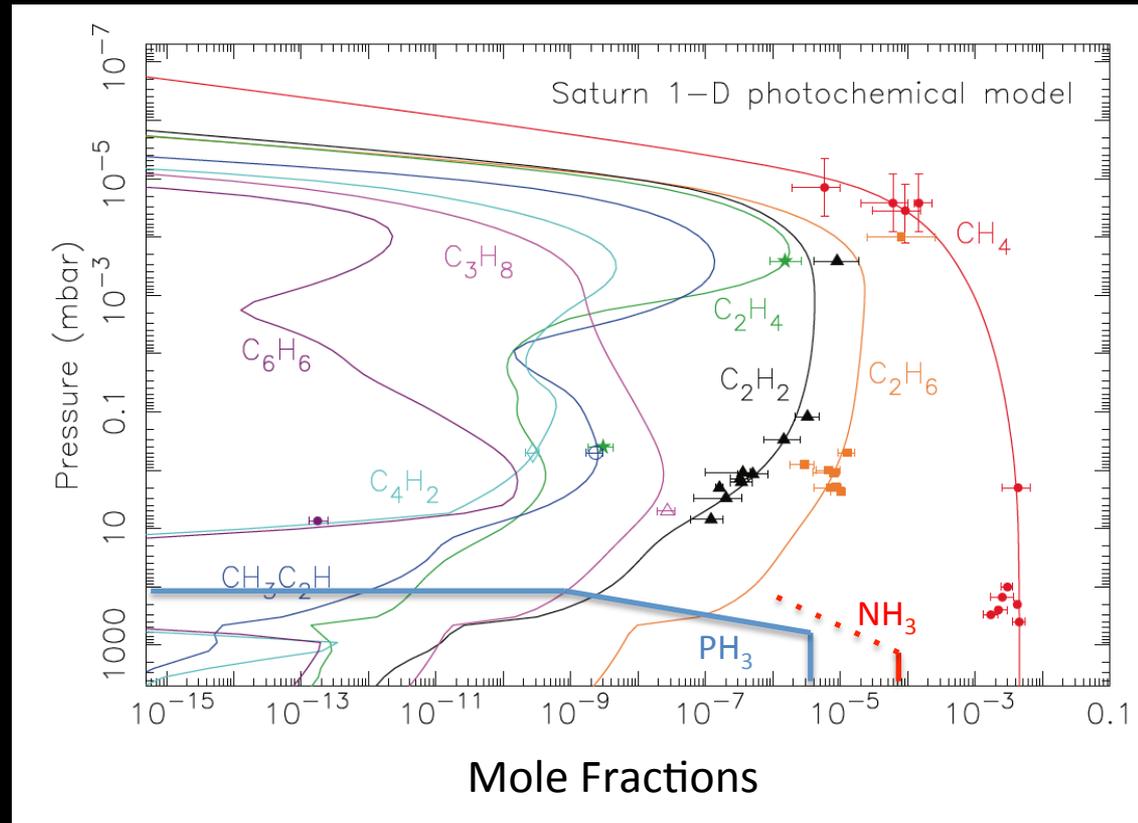


# Part Two: Seasonally-Evolving Composition



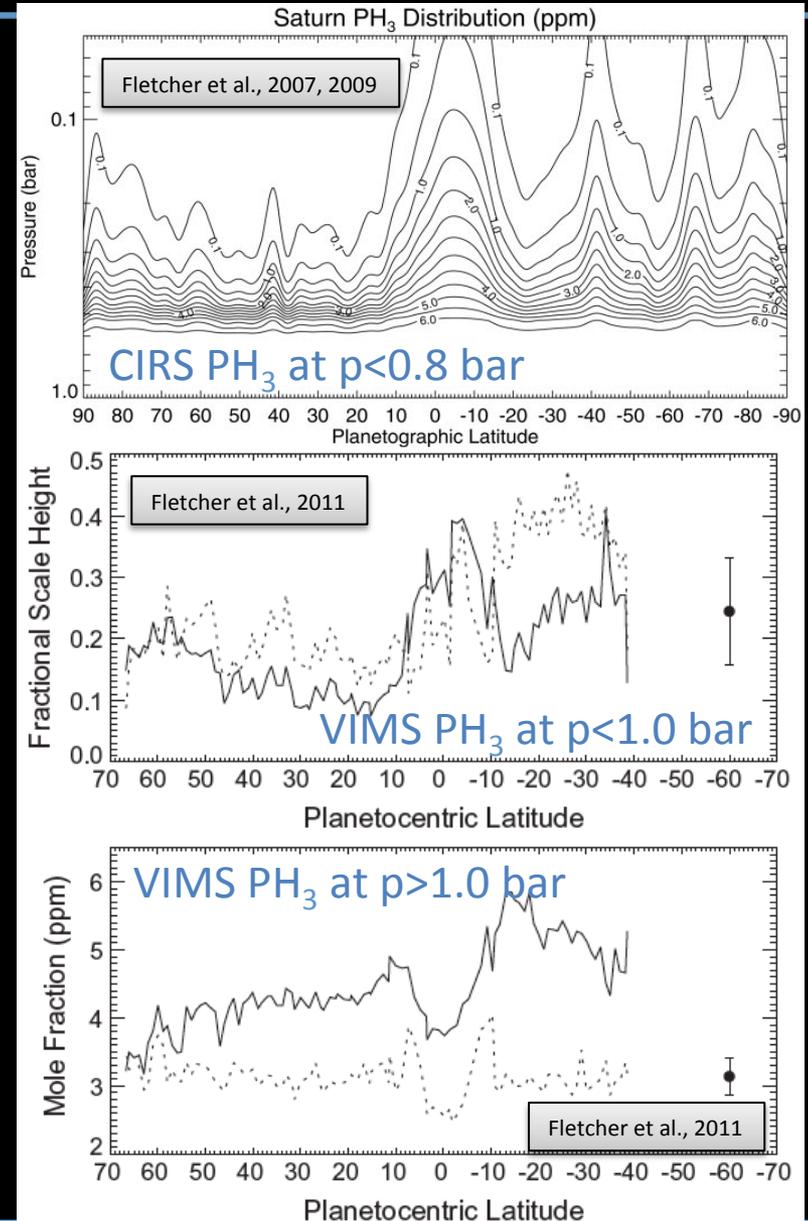
# Introduction: Chemical Soup

- Chemical distributions intricately tied to radiative budget.
- Fouchet et al. (2009) review of all **pre-Cassini measurements**
  - Mostly disc-integrated from the ground and Voyager.
- Cassini **detected asymmetries in stratospheric and tropospheric composition** in southern summer, now hoping to trace reversals.
  - Related to tropospheric/stratospheric circulation and efficiency of photolytic destruction/production.



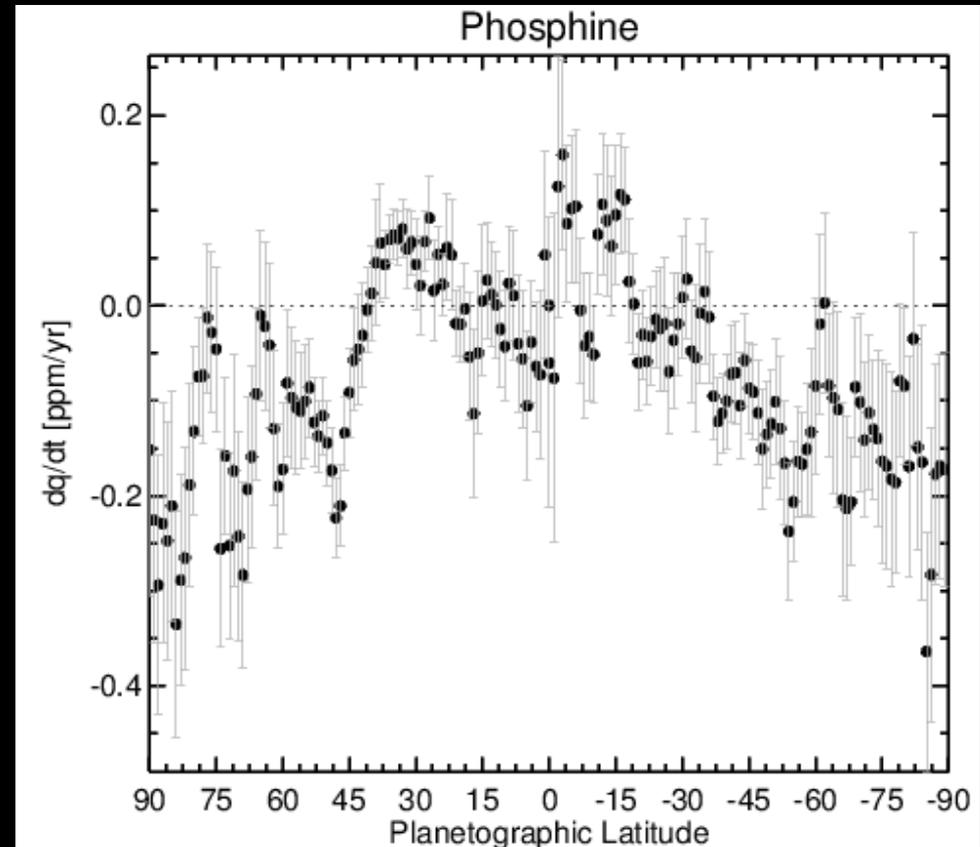
# Phosphine in Southern Summer

- Tropospheric **disequilibrium species**.
  - Well mixed  $p > 500$ -600 mbar.
  - Depleted at  $p < 500$  mbar by photolysis.
- CIRS and VIMS agree on upper tropospheric tropical peak; north-south asymmetry  $p < 1$  bar.
- **Inconsistencies in absolute mixing ratios** still to be resolved.
- More  $\text{PH}_3$  where there's more sunlight?
  - Matches asymmetry in haze opacity, possible shielding effects?
  - More convective mixing in warmer conditions?



# Evolution of Phosphine Distribution

- **PH<sub>3</sub> asymmetry could trace reversal of tropospheric aerosol opacity asymmetry.**
  - Fletcher et al. (2010) reported no significant changes in PH<sub>3</sub> from 2004-2009.
- Latest [preliminary!] results suggest declining abundances in both hemispheres poleward of  $\pm 45^\circ$  – but extremely noisy and hard to explain!
- No changes to equatorial upwelling.

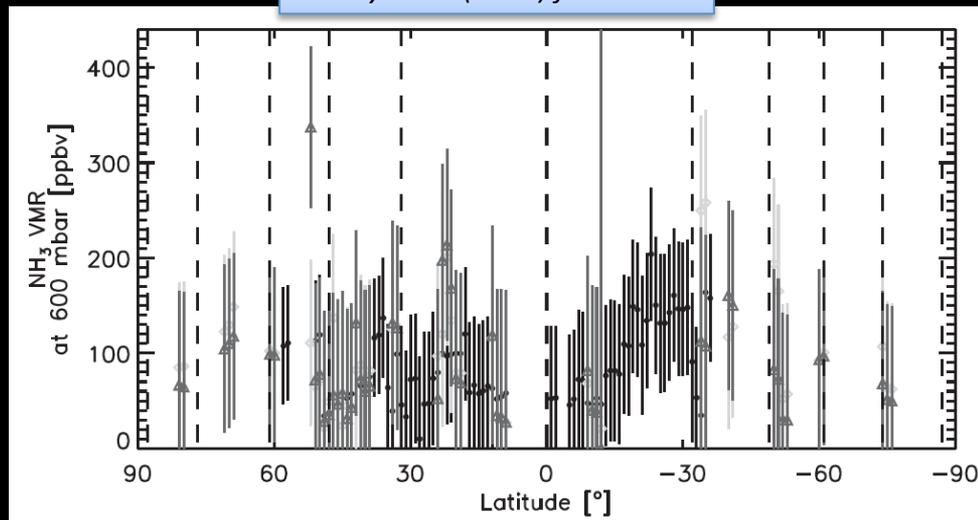


$dq/dt$  at 0.5 bar from ten years of CIRS

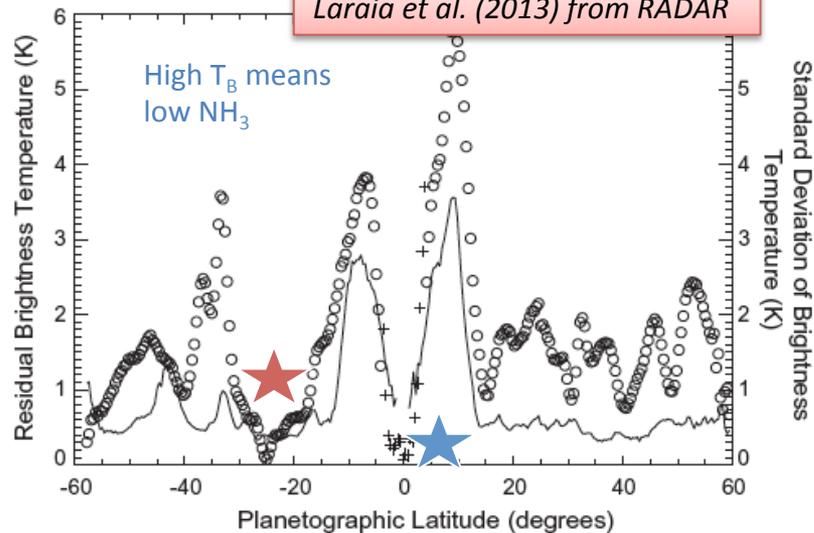
# Unchanging Ammonia?

- Expect  $\text{NH}_3$  photolysis to follow similar trends to  $\text{PH}_3$
- Also expect **more  $\text{NH}_3$  where it's warmer** (saturated abundances increase).
- CIRS observes hints of a north-south asymmetry  $p < 1$  bar.
- VIMS & RADAR see an equatorial maximum between  $\pm 10^\circ$   $p > 1$  bar.

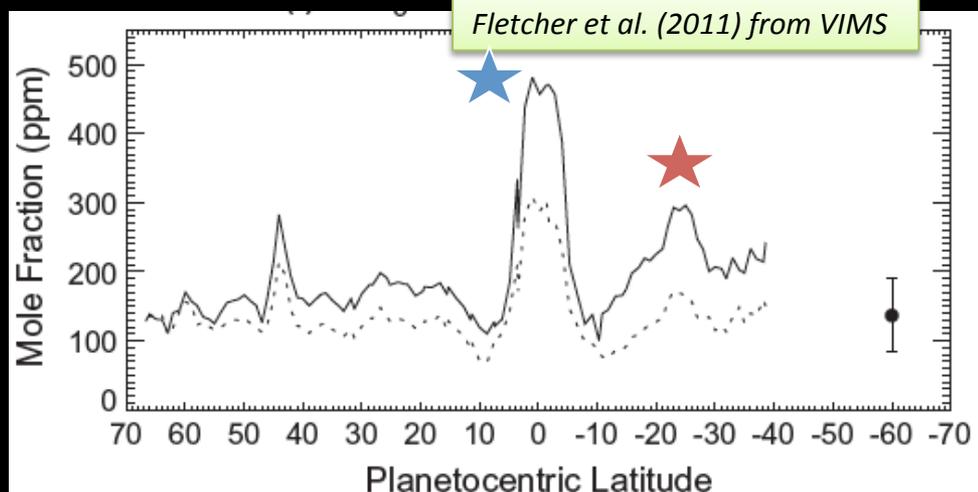
Hurley et al. (2012) from CIRS



Laraia et al. (2013) from RADAR

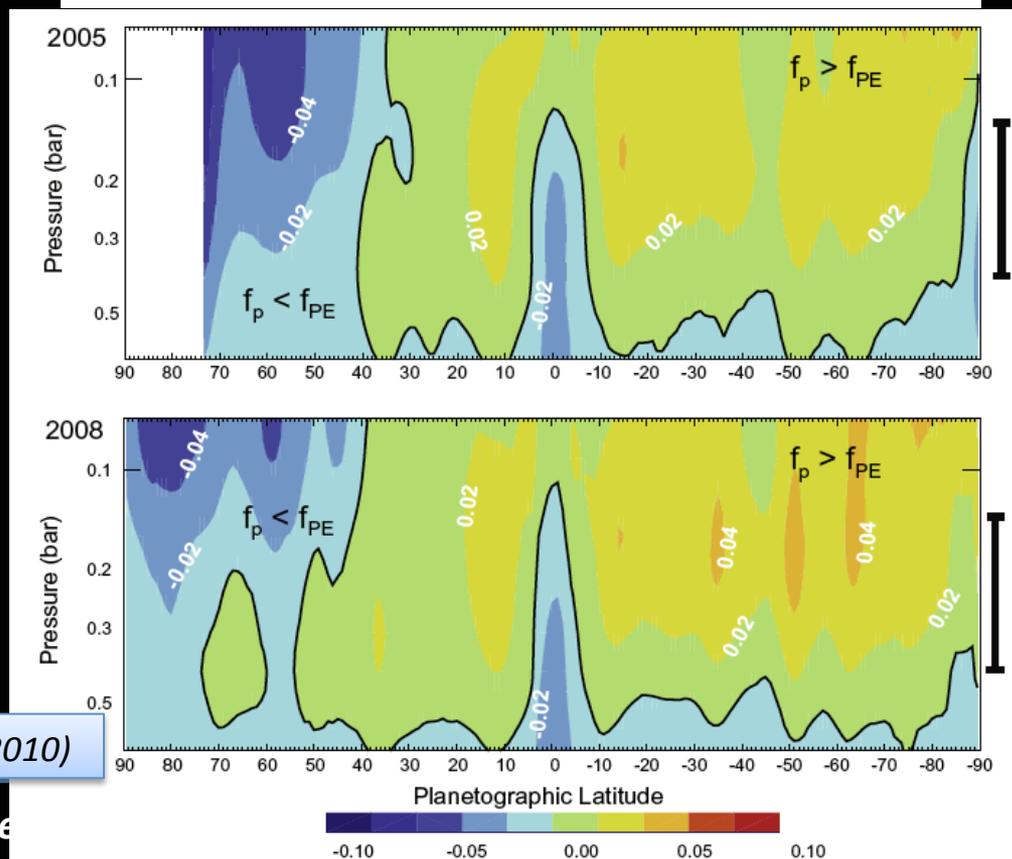
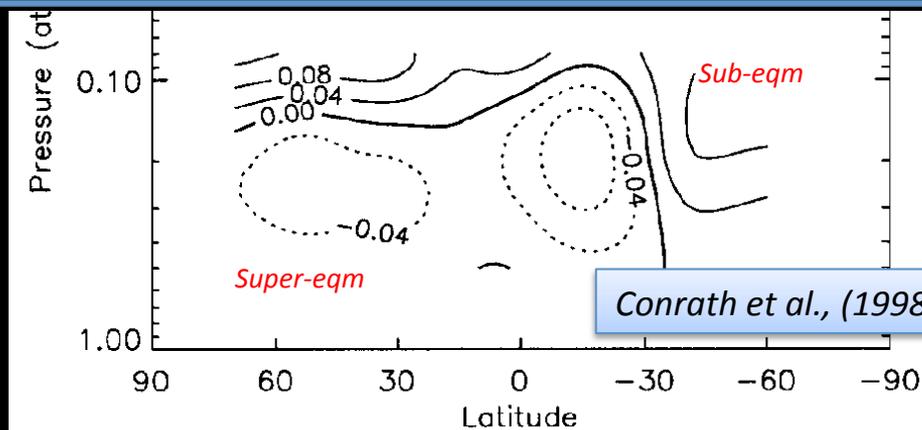


Fletcher et al. (2011) from VIMS

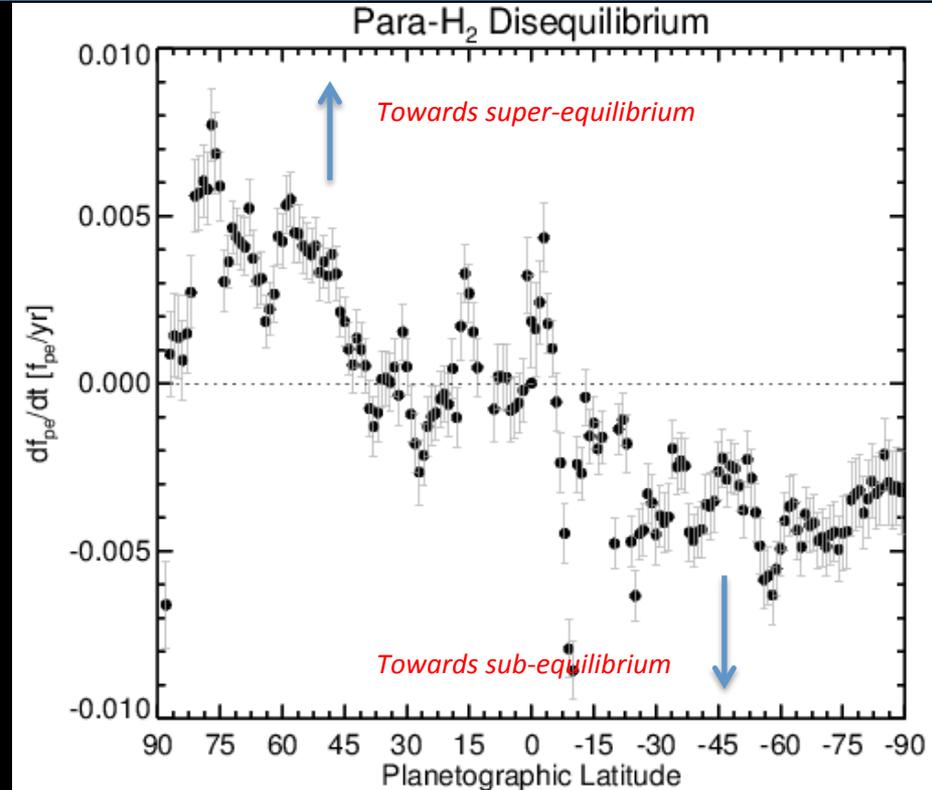
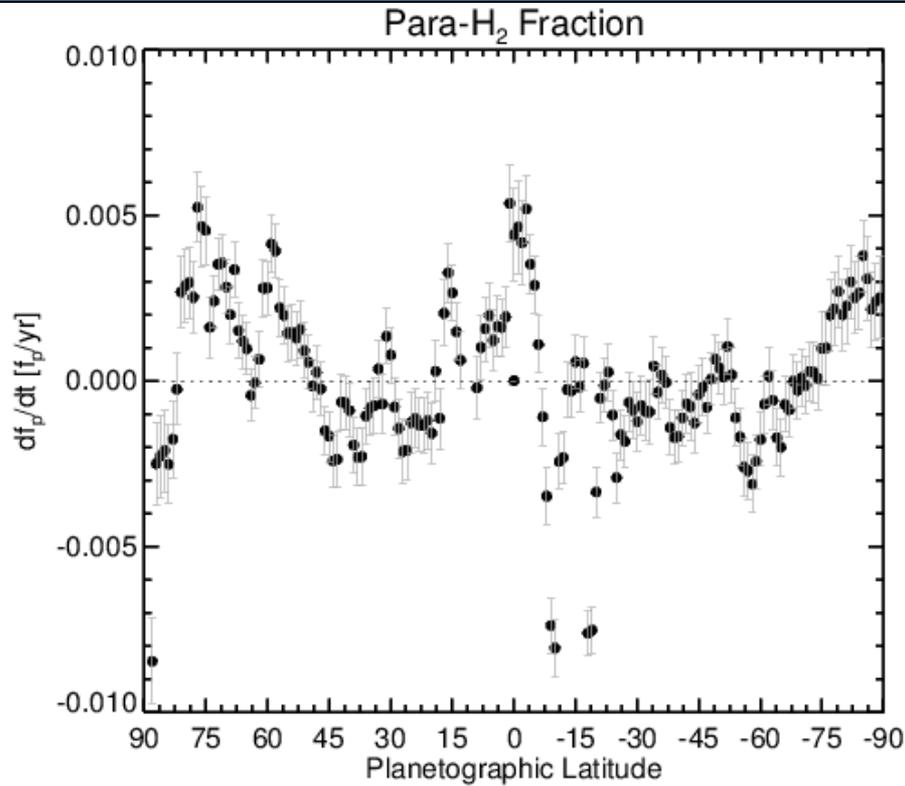


# Para-H<sub>2</sub> Asymmetry

- **Tracer of tropospheric mixing/catalytic efficiency.**
  - Rising air parcels retain low para fraction ( $f_p$ ) of deeper, warmer levels (sub-eqm,  $f_p < f_{eqm}$ ).
  - Sinking parcels have  $f_p > f_{eqm}$  (super-eqm)
- Voyager: Subsidence in the northern spring hemisphere; upwelling in autumn hemisphere.
- Cassini shows different asymmetry to Voyager; equatorial upwelling.



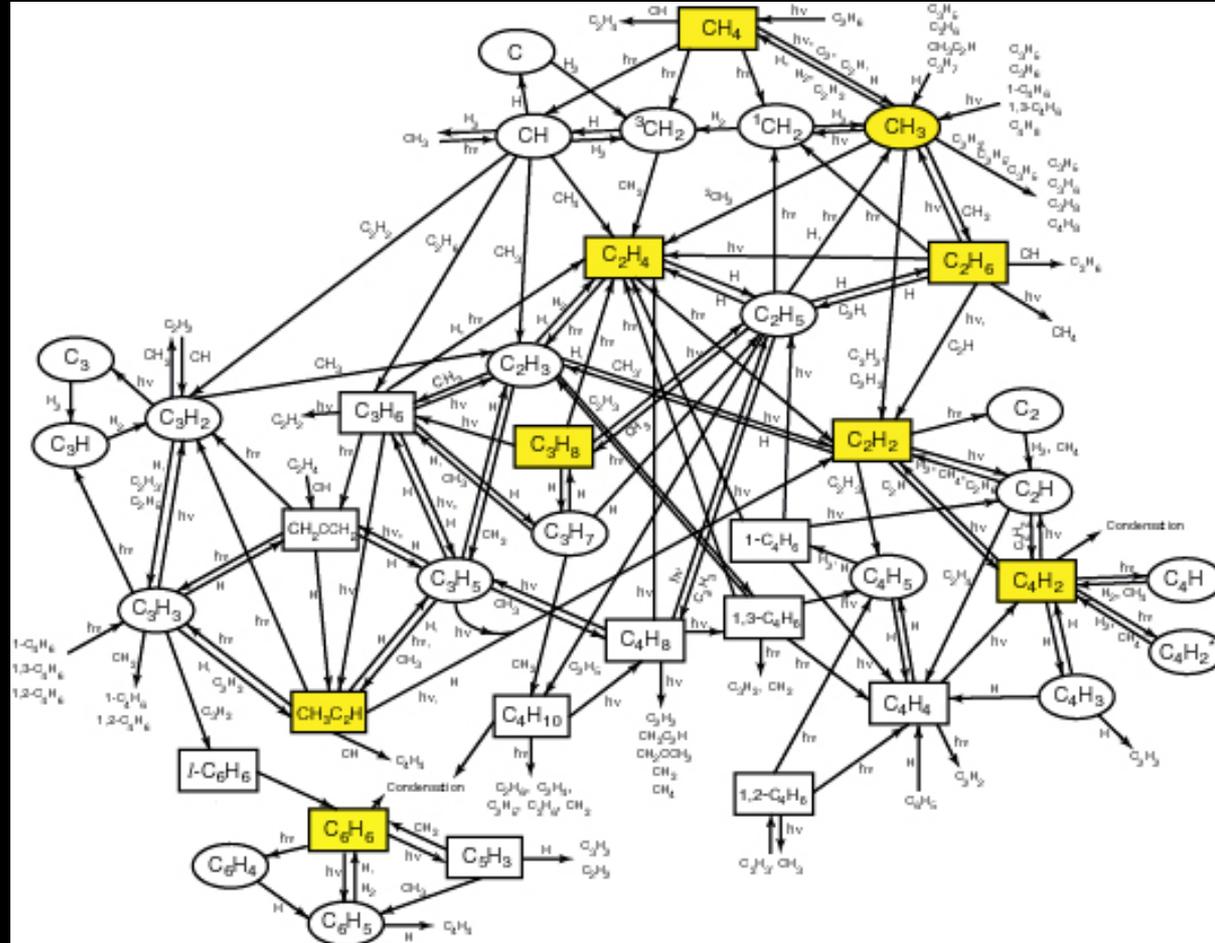
# Latest Para-H<sub>2</sub> Results [Preliminary!]



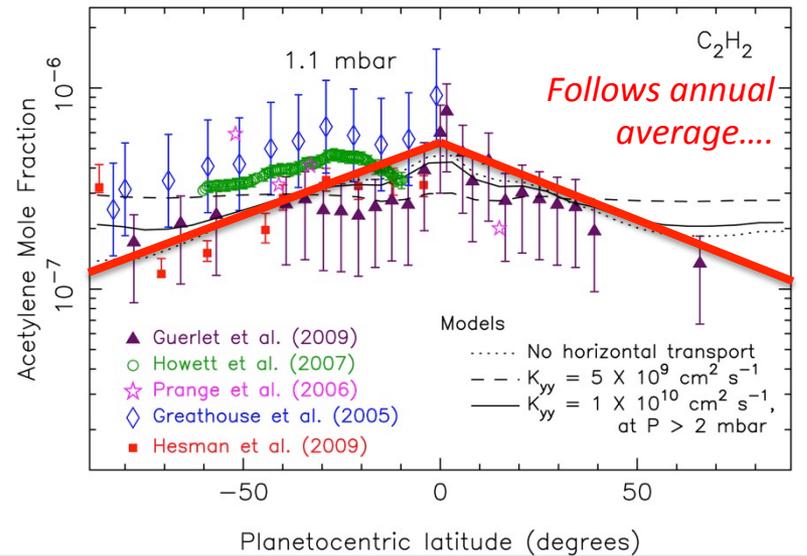
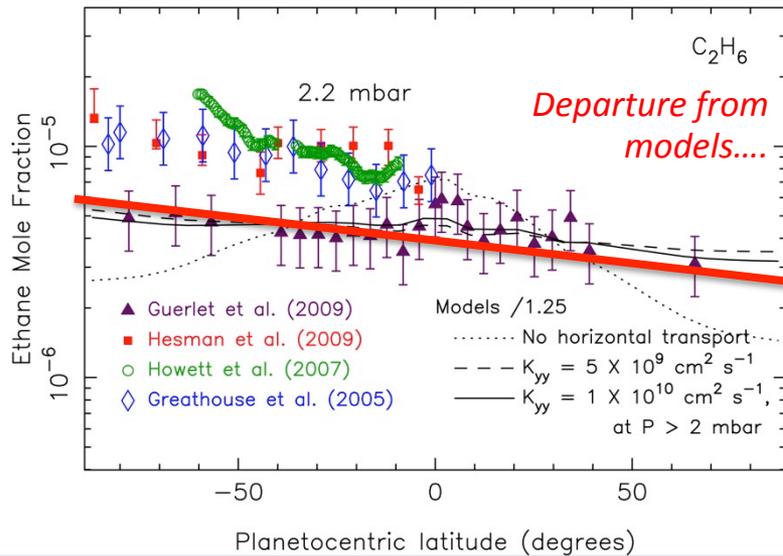
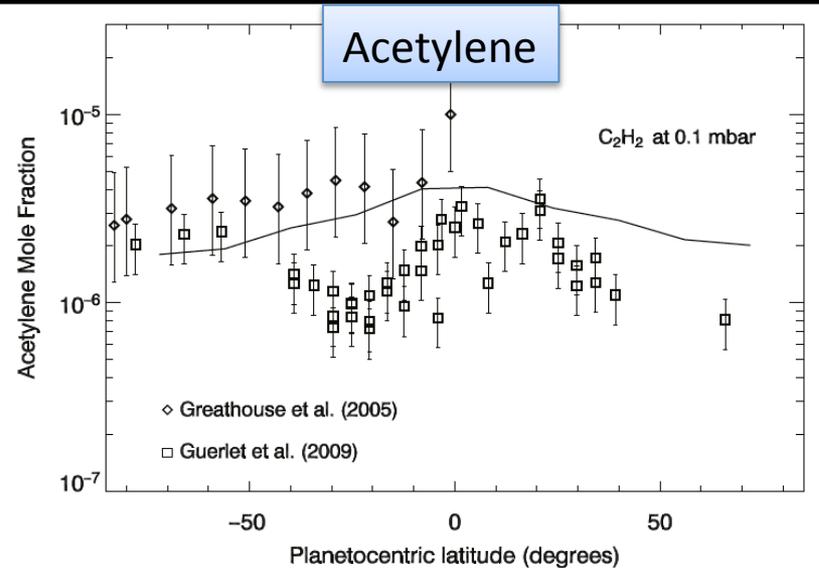
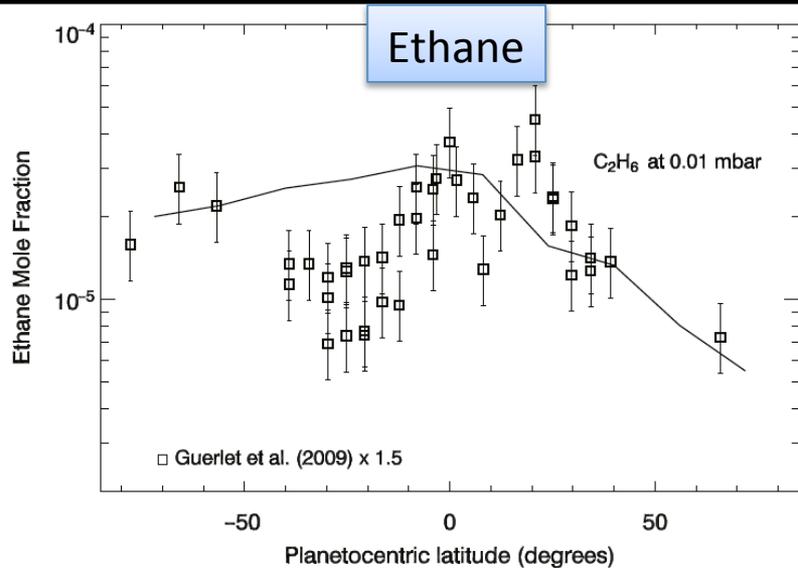
- Para-H<sub>2</sub> has not changed significantly, **disequilibrium asymmetry has started to reverse** (slowly) due to equilibrium para-H<sub>2</sub> changes.
  - Equilibration timescale  $\sim$  century @ 200 mbar, **don't expect big changes**.
- Northern latitudes have become super-equilibrium, southern latitudes tending to sub-equilibrium, tracking temperature/aerosol asymmetry reversal.

# Stratospheric Hydrocarbons

- Before Cassini, 1D chemistry-diffusion models limited to fixed seasons & few latitudes.
- Full seasonal 1D (Kzz) diffusive photochemistry model (Moses and Greathouse, 2005) predict:
  - Long-lived CxHy more affected by transport and follow annual average insolation ( $C_2H_6$  and  $C_3H_8$ ).
  - Short-lived CxHy continue to show seasonal asymmetries at  $p > 1$  mbar ( $CH_3$ ,  $C_2H_2$ ,  $C_2H_4$ ,  $CH_3C_2H$ , and  $C_4H_2$ )
  - Species should follow annual-average insolation at  $p > 1$  mbar
- No published 2D models (i.e., Kzz and Kyy mixing) to date.
  - *Work in progress....*



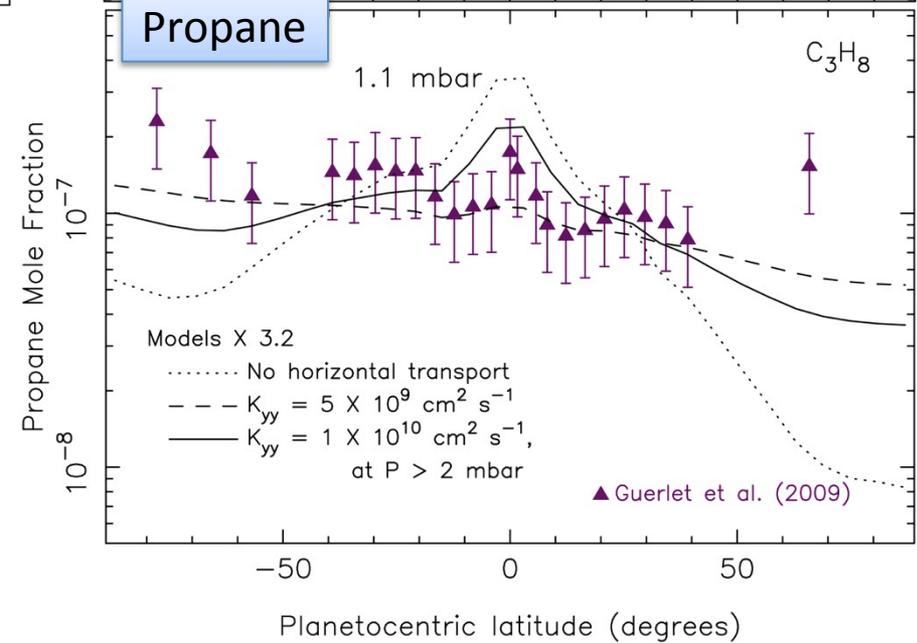
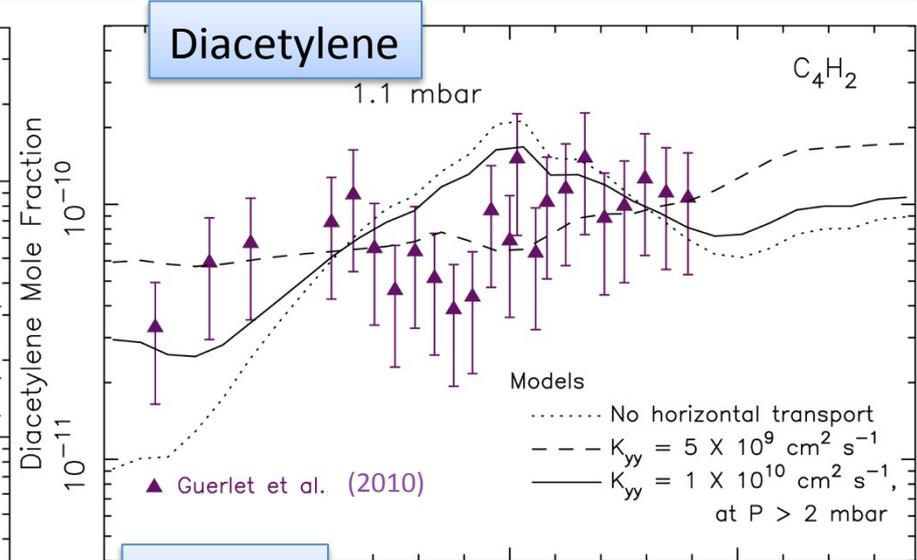
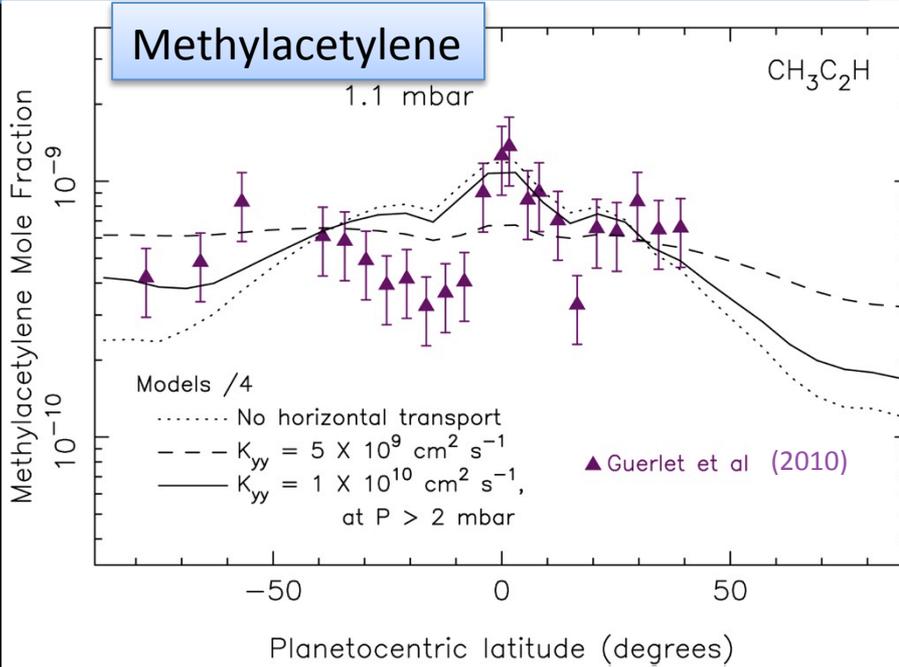
# C<sub>2</sub>H<sub>2</sub> & C<sub>2</sub>H<sub>6</sub> Summer Observations



Greathouse et al. (2005), Howett et al. (2007), Hesman et al. (2009), Guerlet et al. (2009), Sinclair et al. (2013)



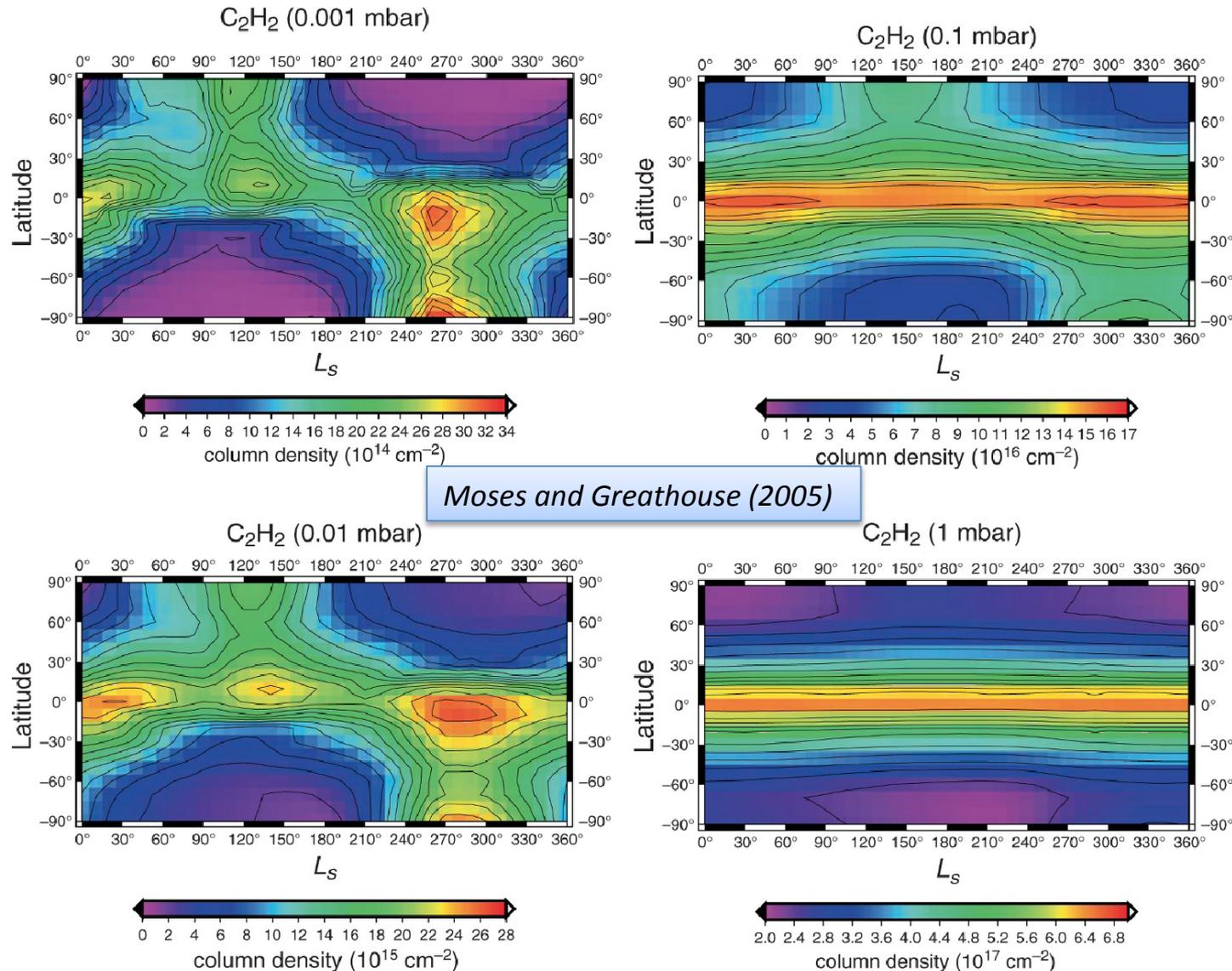
# C<sub>3</sub>H<sub>x</sub> and C<sub>4</sub>H<sub>x</sub> Summer Observations



- Asymmetries in C<sub>3</sub>H<sub>4</sub> and C<sub>4</sub>H<sub>2</sub> not captured by models, higher in mid-northern latitudes.
- C<sub>3</sub>H<sub>8</sub> relatively flat, hints of elevation in south.
- Departures may be caused by transport.

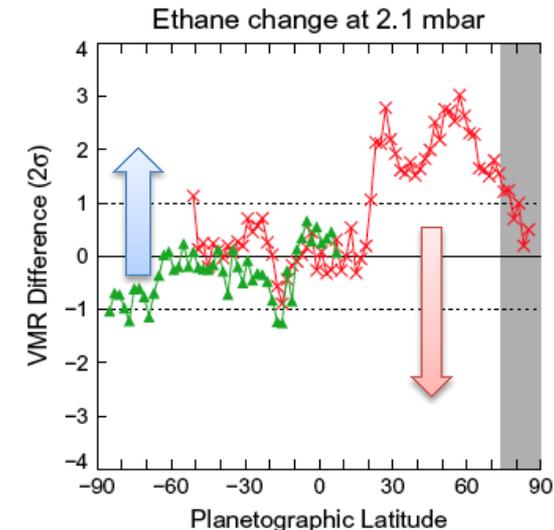
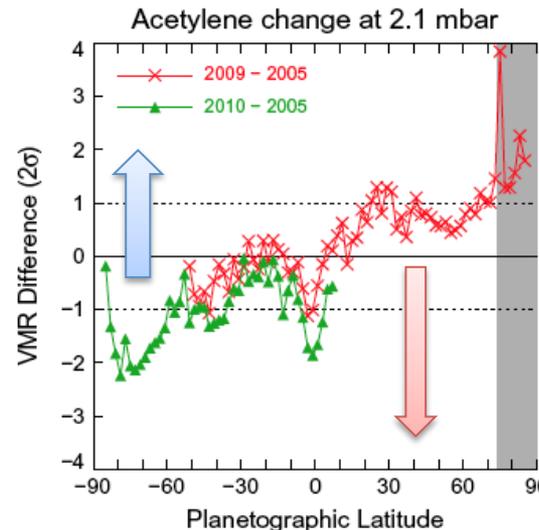
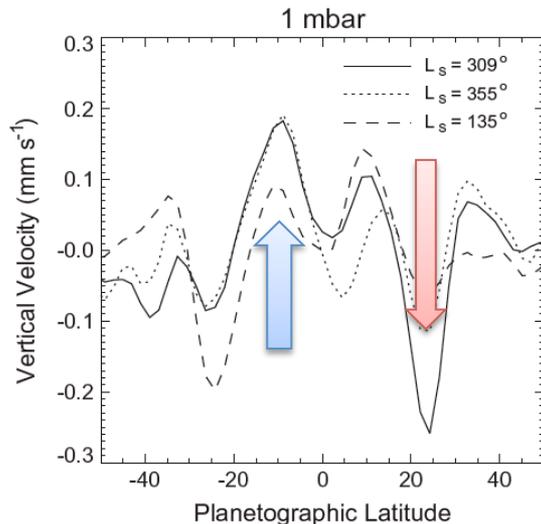
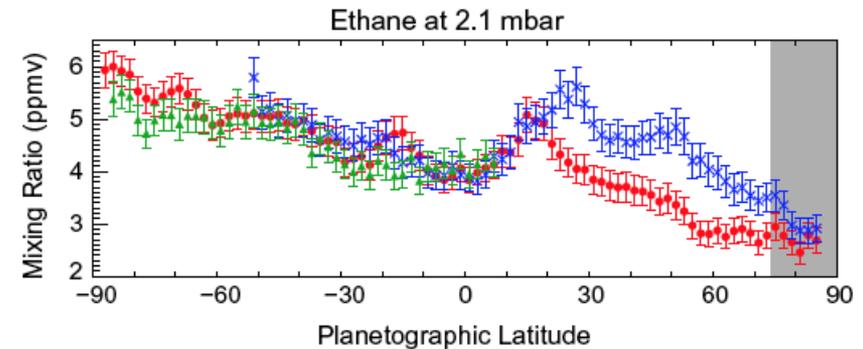
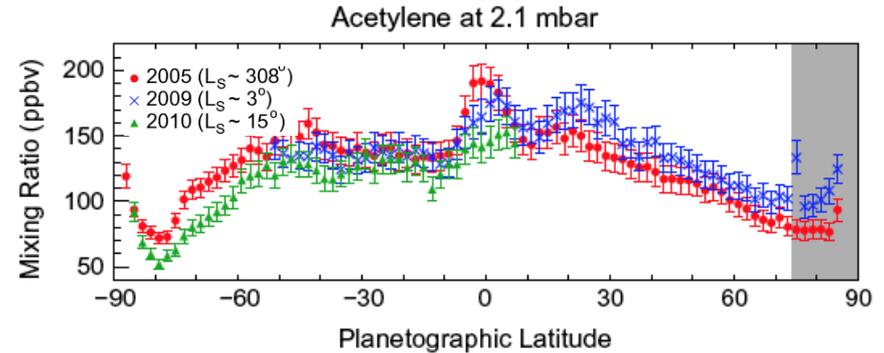
# Temporal CxHy I: Predictions

- Summertime pulse moves downwards with phase lag.
- Strong asymmetries expected at microbar pressures.
- Negligible asymmetries at millibar pressures.
- Introduction of Kyy may improve reproduction of data.



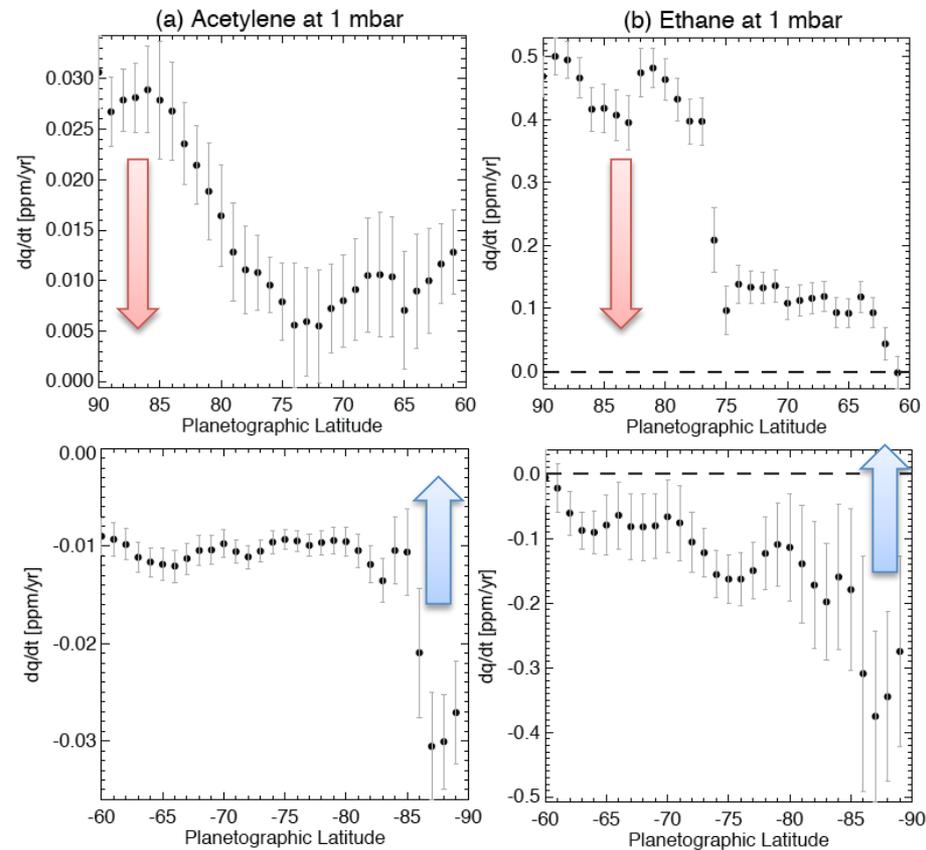
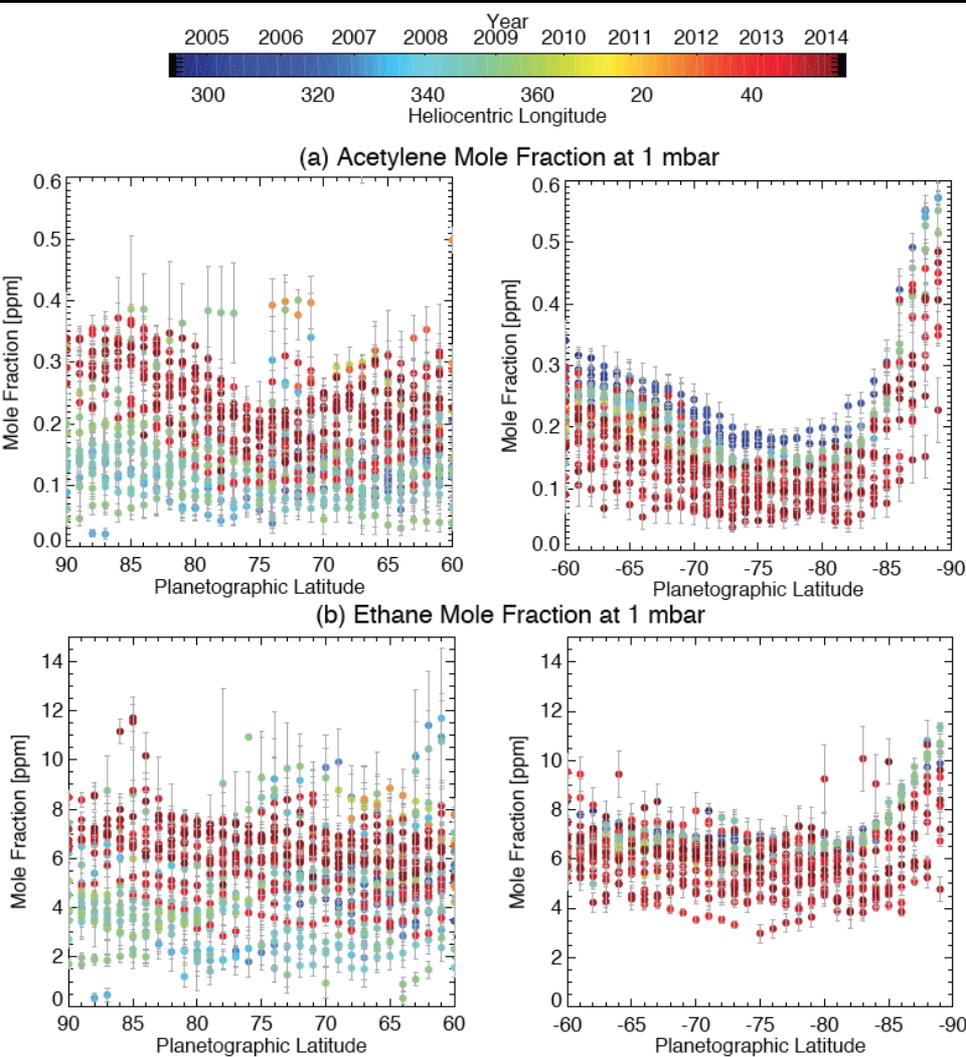
# Temporal CxHy II: Global Trends

- Sinclair et al. (2013) first seasonal study of  $C_2H_2$  and  $C_2H_6$ .
  - Increasing abundances in northern spring; decreasing in southern autumn.
- Low-latitude changes consistent with simulation of Friedson and Moses (2012).
- Correspondence at high latitudes unclear.



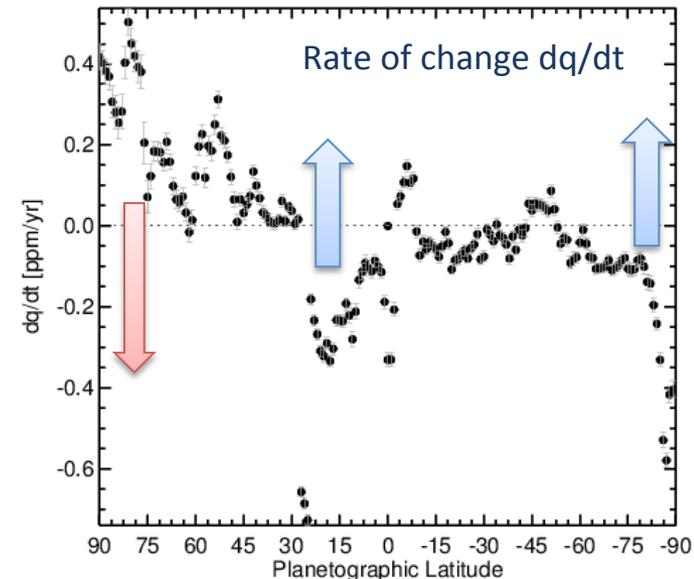
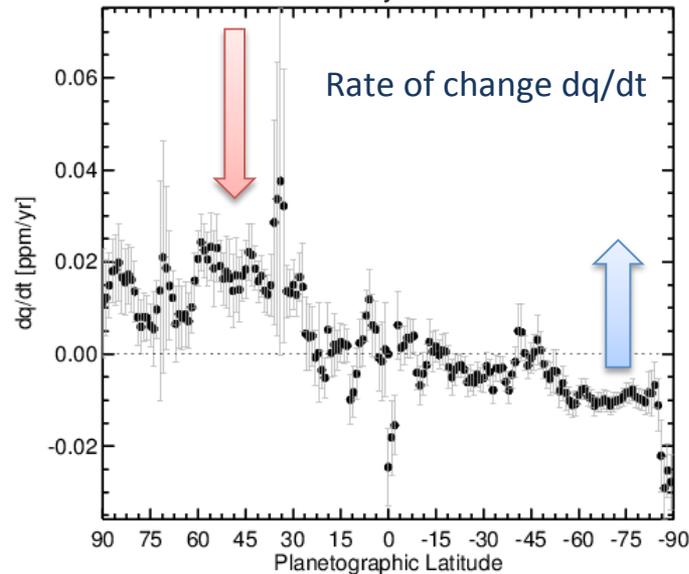
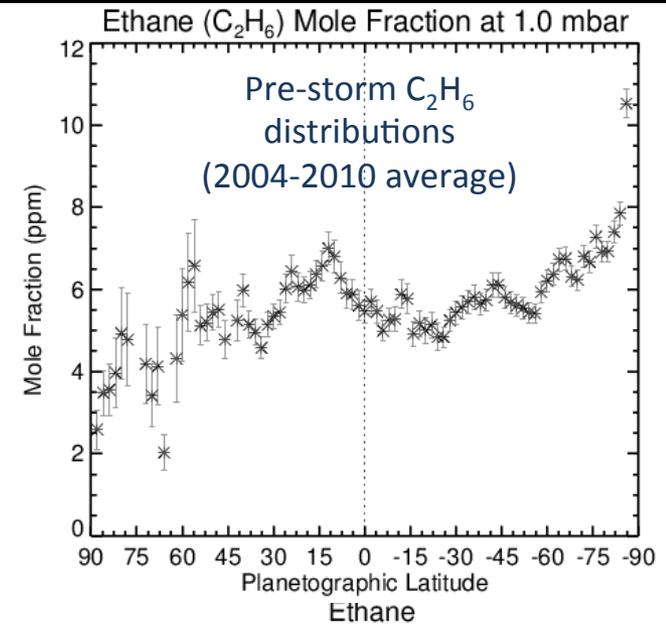
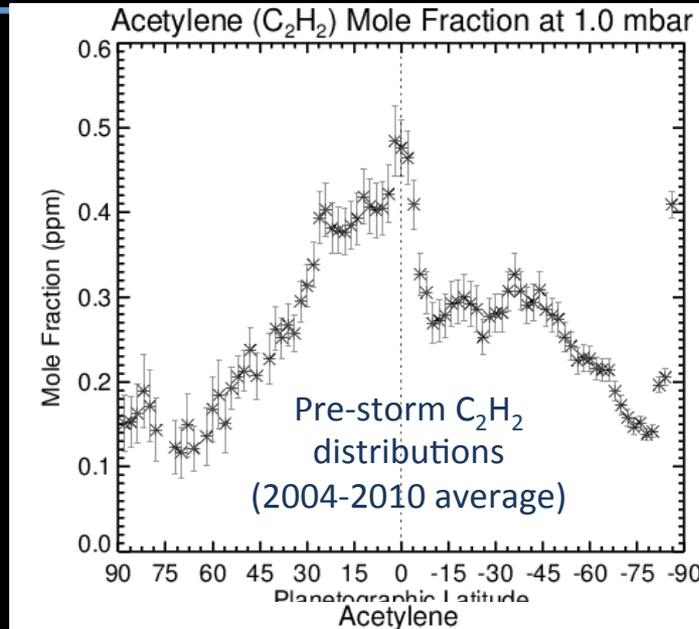
# Temporal CxHy III: Polar Variability

- $C_2H_2$  and  $C_2H_6$  changes most dramatic within polar vortices:
  - $dq/dt$  measured over ten years.
  - Continuity implies  $|w| \sim 0.1$  mm/s



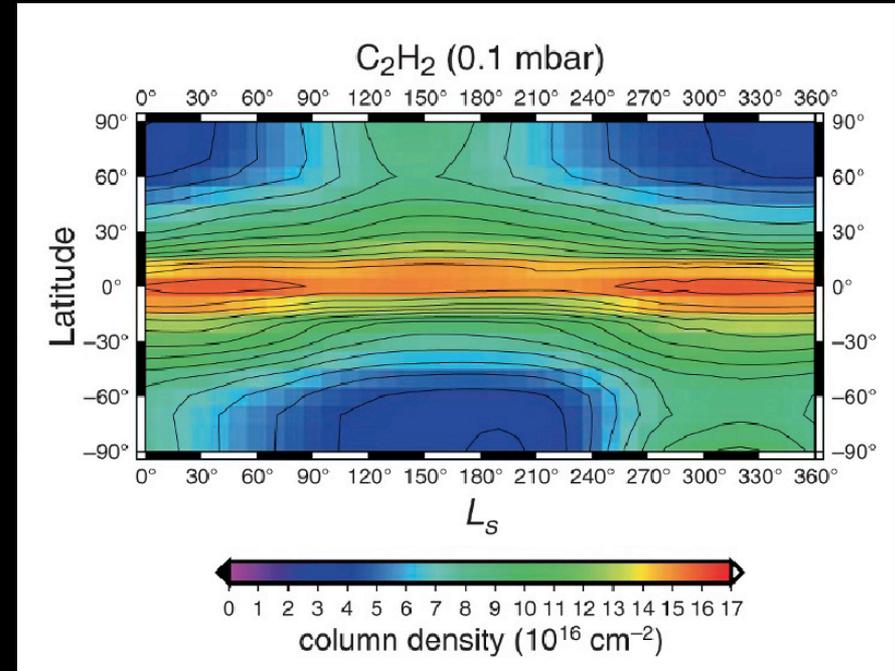
# Temporal CxHy IV: Latest Results

- Preliminary investigation of CIRS data suggest ethane and acetylene asymmetries are starting to reverse at 1 mbar.
- Hemispheric trends perturbed by tropical and polar circulation cells.



# Summary of Part Two: Composition

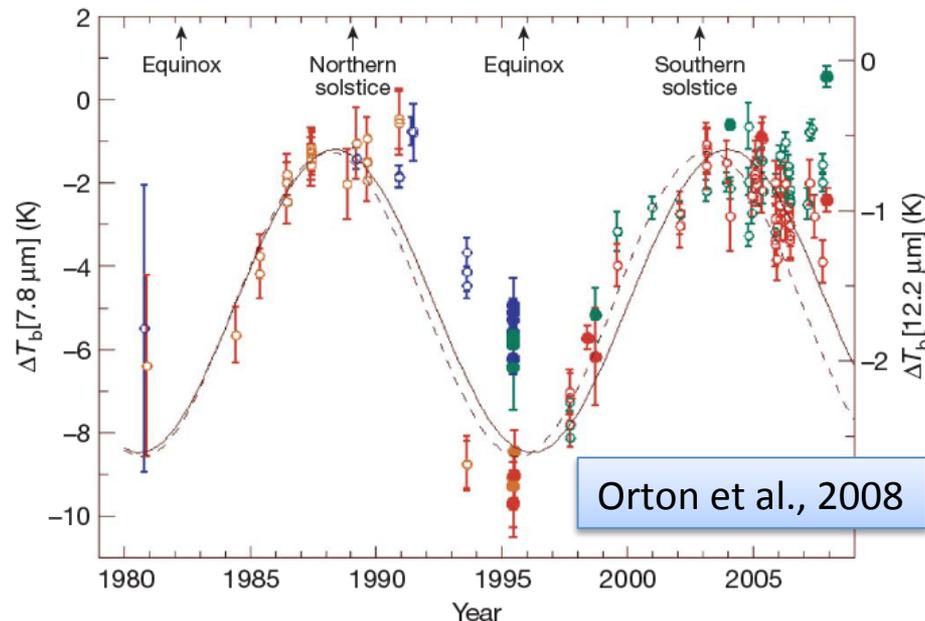
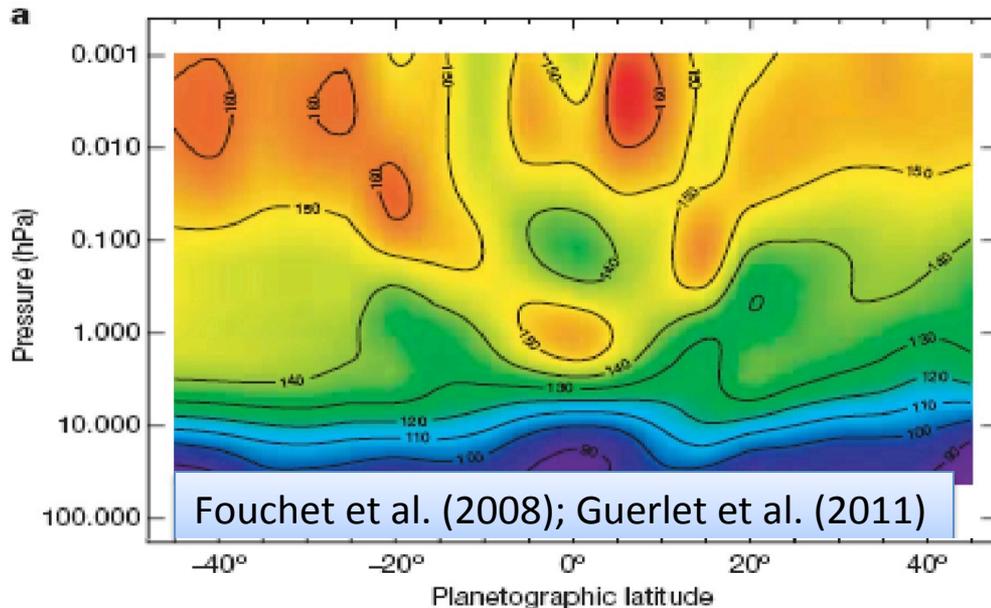
1. Tropospheric and stratospheric species influenced by photochemistry/circulation all show **summertime asymmetries**.
2. Diffusive **photochemistry models can reproduce some** (but not all) observed trends.
3. Latest Cassini results show **asymmetries reversing post-equinox**.
4. Disentangling chemical and transport effects means use of species as tracers isn't always clear-cut.



# Briefly.... Seasonally Modulated Phenomena

i.e., Other topics we'll touch on in Chapter 10

# Wave Phenomena

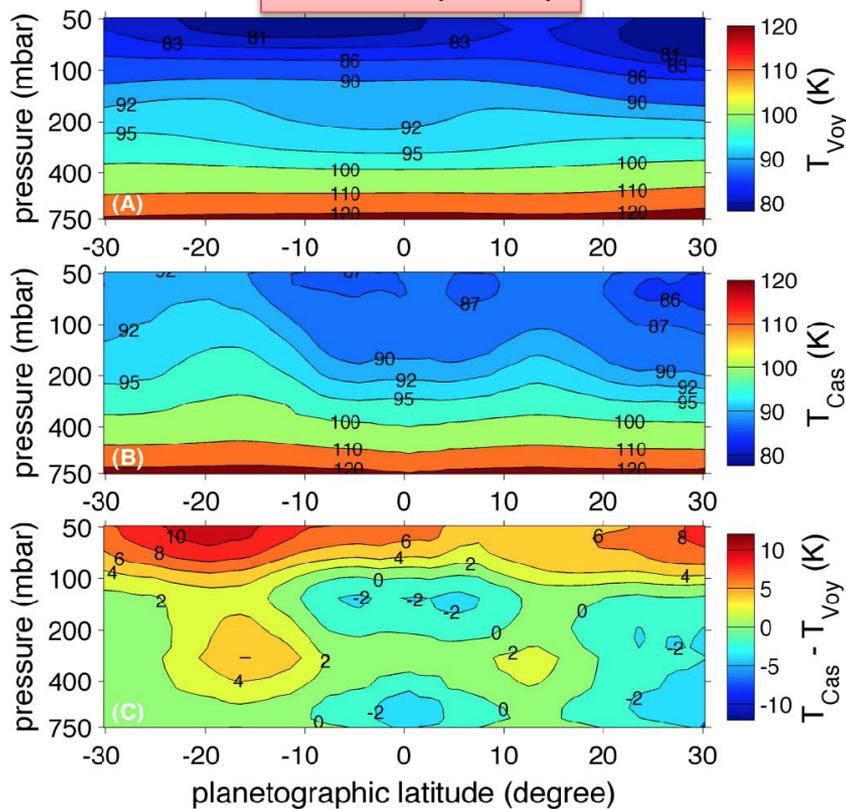


- **Tropical stratosphere dominated by effects of quasi-periodic oscillation**
  - (Fouchet et al. 2008; Orton et al., 2008; Fletcher et al., 2010; Schinder et al., 2011; Guerlet et al., 2011).
- Downward motion of a chain of warm/cool airmasses.
- Upward-propagating planetary waves interacting with the mean zonal flow.
- Phase of SSAO connected to seasonally-reversing Hadley circulation? [Friedson & Moses, 2012]
- **Recent disruption of SSAO due to storm beacon?**

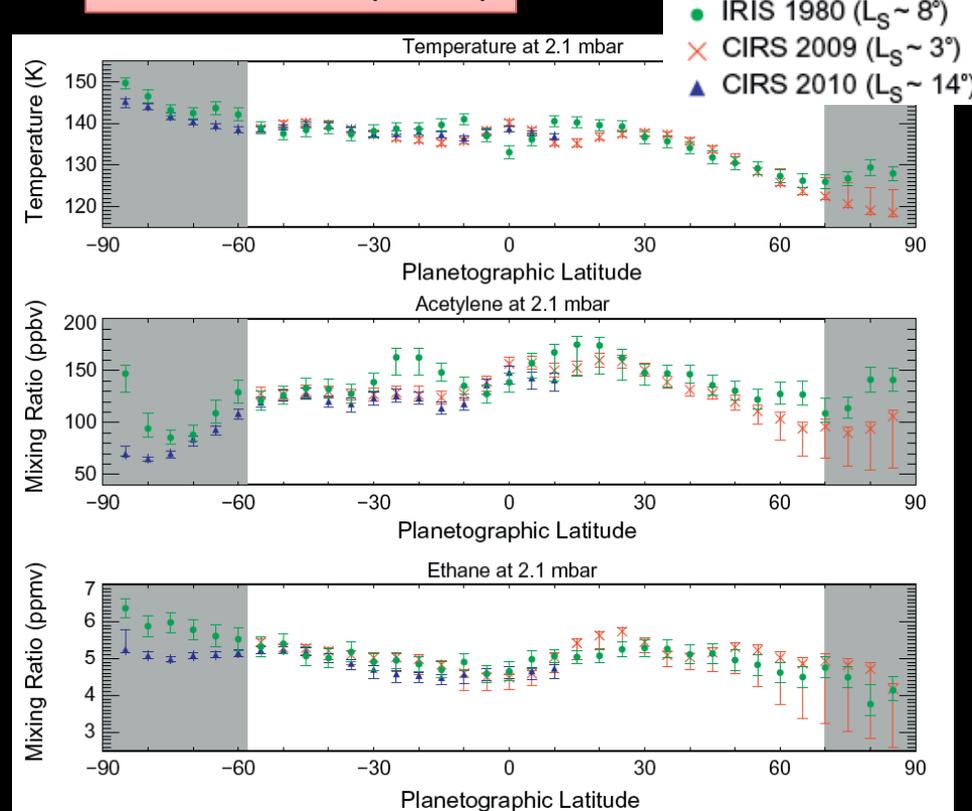
# Interannual Variability?

- Tropospheric [Li et al., 2013] and stratospheric [Sinclair et al., 2014] conditions different in Voyager/Cassini epochs despite same  $L_s$ .
- Viewing **a different phase of SSAO**, not quite semi-annual?

Li et al. (2013)



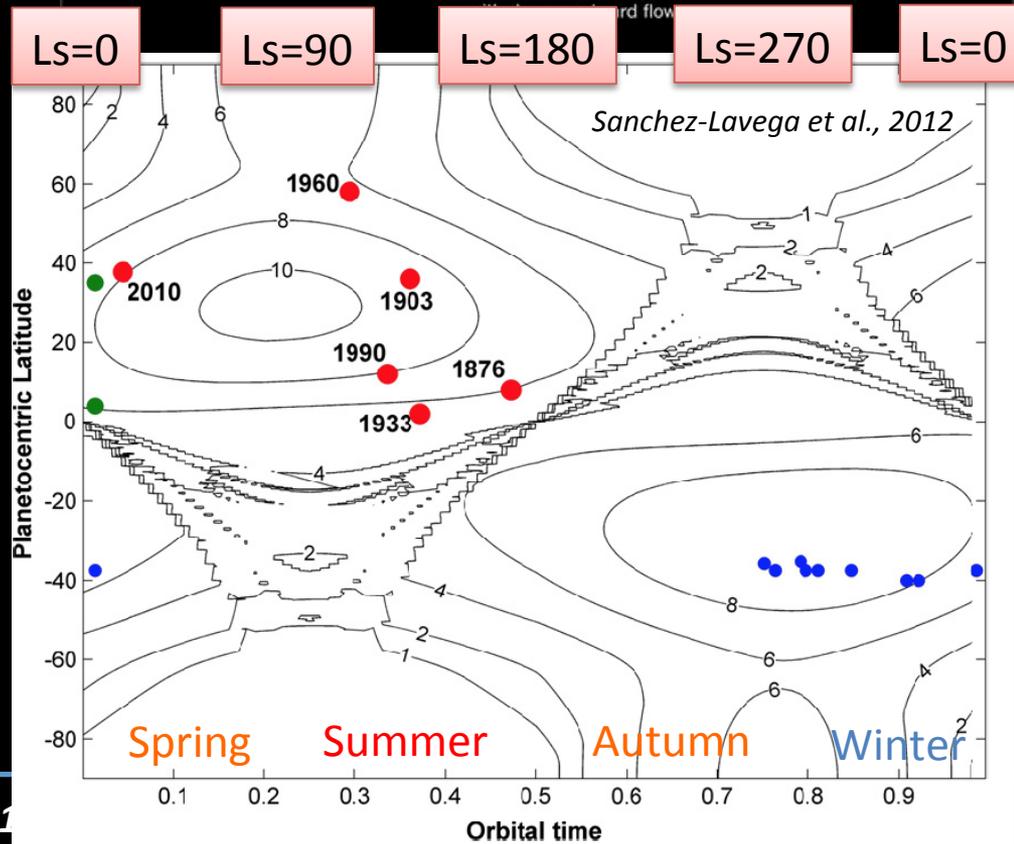
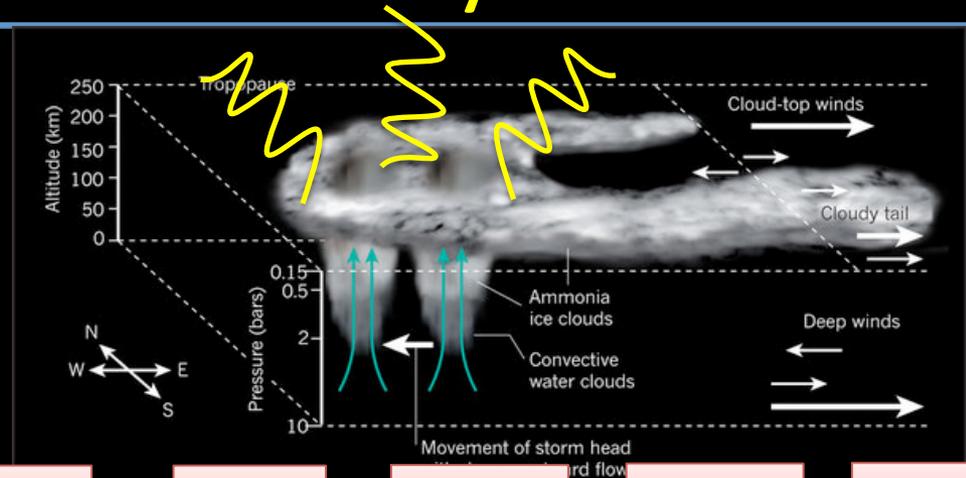
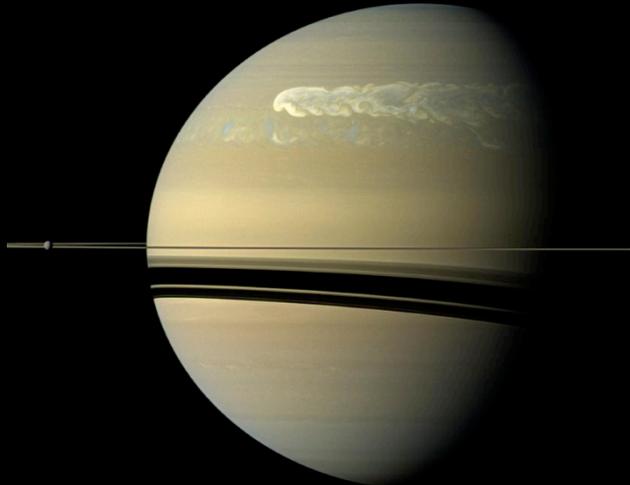
Sinclair et al. (2014)



# Annual Storm Activity

- Convective storms and vortices localised in 'storm alleys'.
  - Intensity/frequency decreasing in southern autumn, increasing in northern spring?
- Seasonal modulation of GWS phenomenon despite eruption from great depth
  - Modification of stability of background troposphere.

Feb 25, 2011



# Conclusions

- Saturn's seasons have been **tracked for more than 30 years**, including 10 years of continuous orbital reconnaissance.
- **Temperature asymmetries** discovered in the 1970s; tracked globally by Cassini:
  - Follows seasonal radiative predictions except in regions of strong dynamic activity (equator, poles).
  - Implications for seasonally reversing mid-atmosphere winds and circulation.
- **Compositional asymmetries** discovered by Cassini
  - Tropospheric  $\text{PH}_3$  and para- $\text{H}_2$
  - Stratospheric hydrocarbons.
  - Only partially explained by photochemical models lacking comprehensive transport schemes.
  - Reversing asymmetries post-equinox.

- **By northern summer:**
  - Stratosphere temperature asymmetry reversed, but not as warm as southern summer.
  - $\text{PH}_3$ , para- $\text{H}_2$  and  $\text{CxHy}$  asymmetries reversed or reversing.
  - South polar vortex dissipated; strong methane and hydrocarbon emission from new north polar vortex.
- *Watch this space...*